The black hole - host galaxy relation for very low-mass quasars J. Sanghvi<sup>1</sup>, J.K. Kotilainen<sup>2</sup>, R. Falomo<sup>3</sup>, R. Decarli<sup>4</sup>, K. Karhunen<sup>1</sup>, M. Uslenghi<sup>5</sup>

## Abstract

We have investigated the  $M_{BH}$  -  $M_{host}$  log-linear relation for a sample of 37 quasars with low black hole masses ( $10^7 M_{\odot} < M_{BH} < 10^{8.3} M_{\odot}$ ) at 0.5 < z < 1.0. For 25 quasars, we detected the presence of the host galaxy from deep near-infrared H-band imaging, whereas upper limits for the host galaxy luminosity (mass) were estimated for the 12 unresolved quasars. 75% of quasars were disc dominated. We advocate secular evolution of discs of galaxies being responsible for the relatively strong disc domination.

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