

The systematic search for $z \gtrsim 5$ active galactic nuclei in the *Chandra* Deep Field South

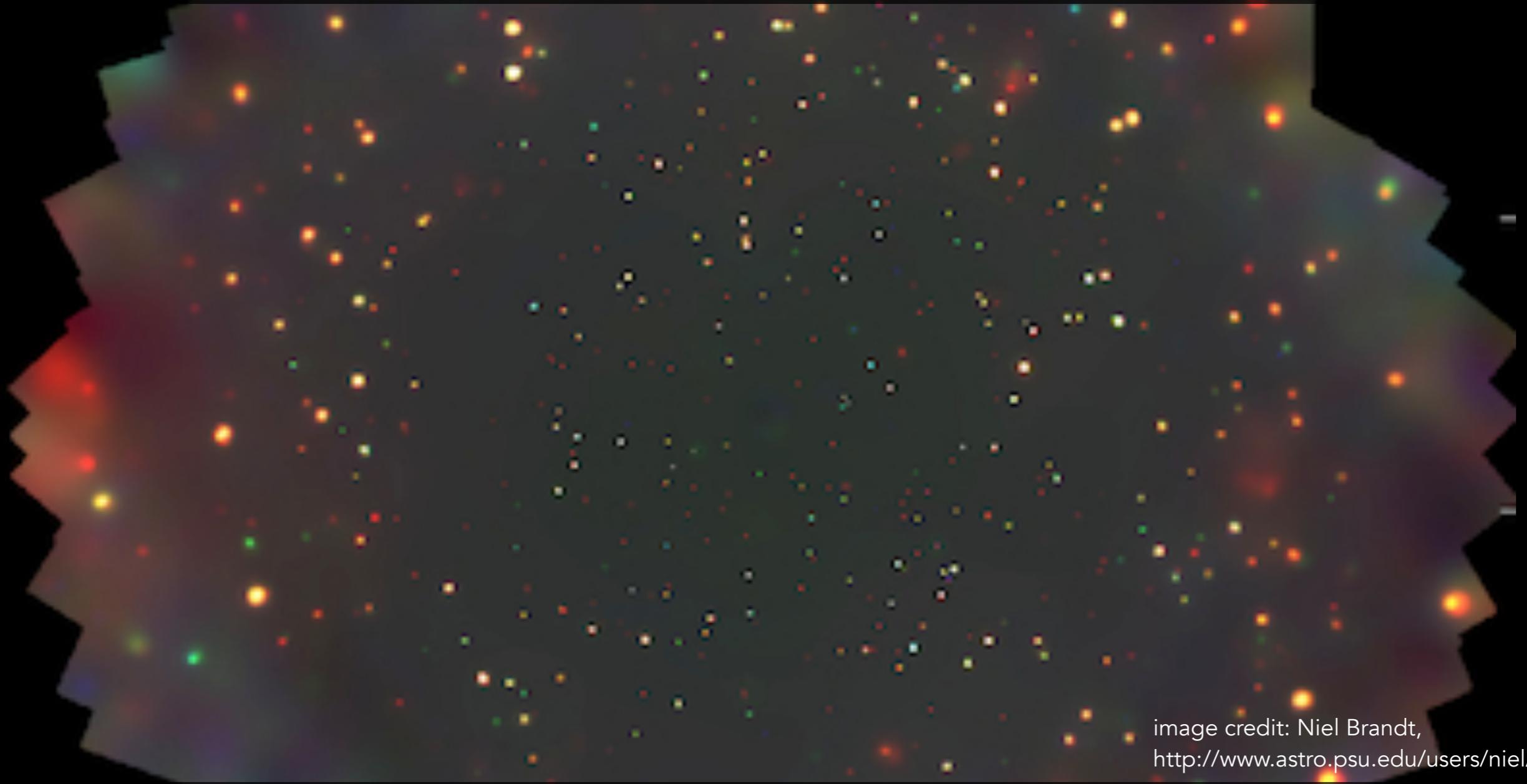


image credit: Niel Brandt,
<http://www.astro.psu.edu/users/niel/cdfs/>

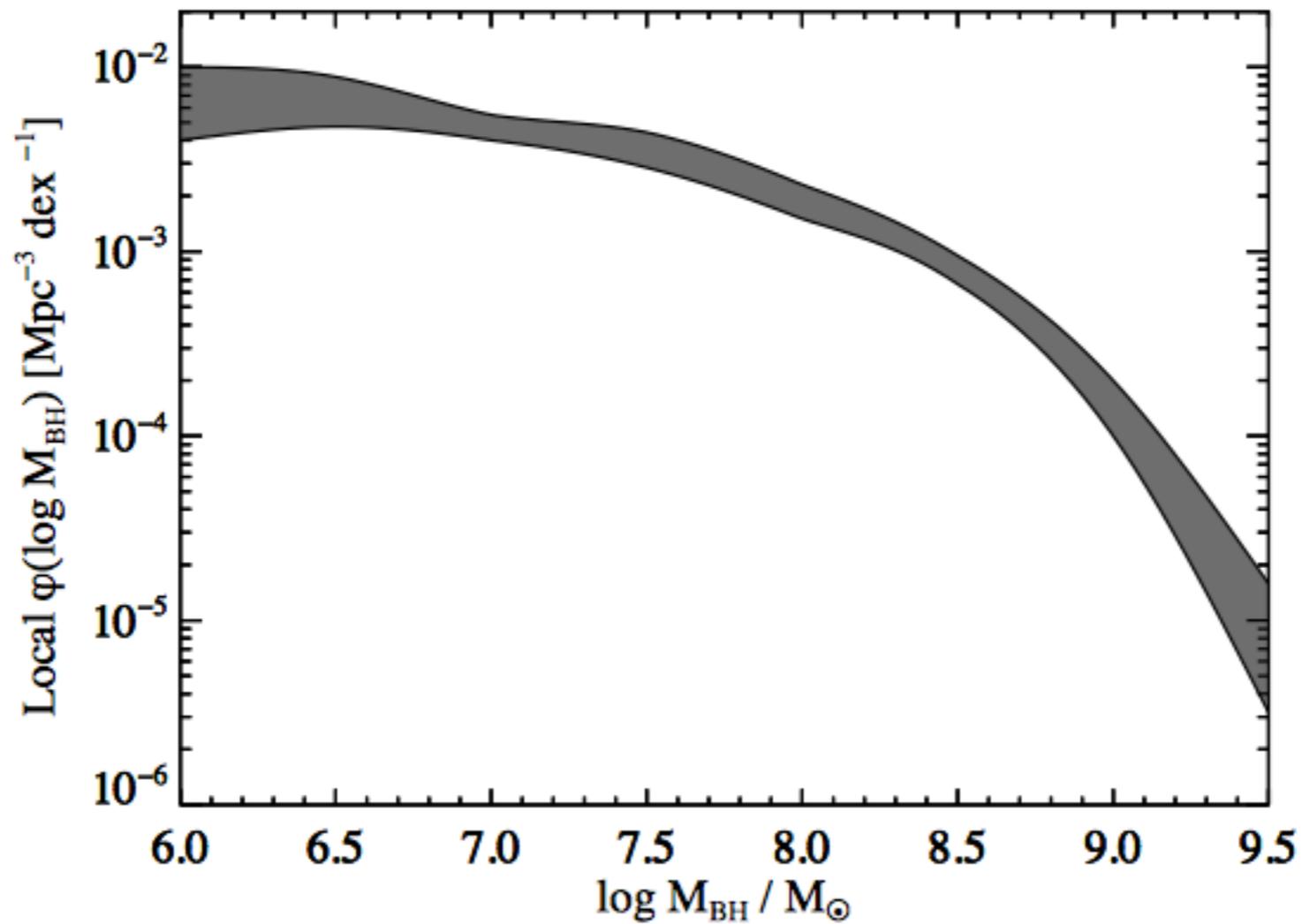
Anna Weigel, ETH Zurich

Kevin Schawinski, Ezequiel Treister, Meg Urry, Michael Koss, Benny Trakhtenbrot

arxiv:1501.06580, MNRAS in press

Motivation

Local BH mass function:



Kelly & Merloni, 2011

Motivation

Local BH mass function:

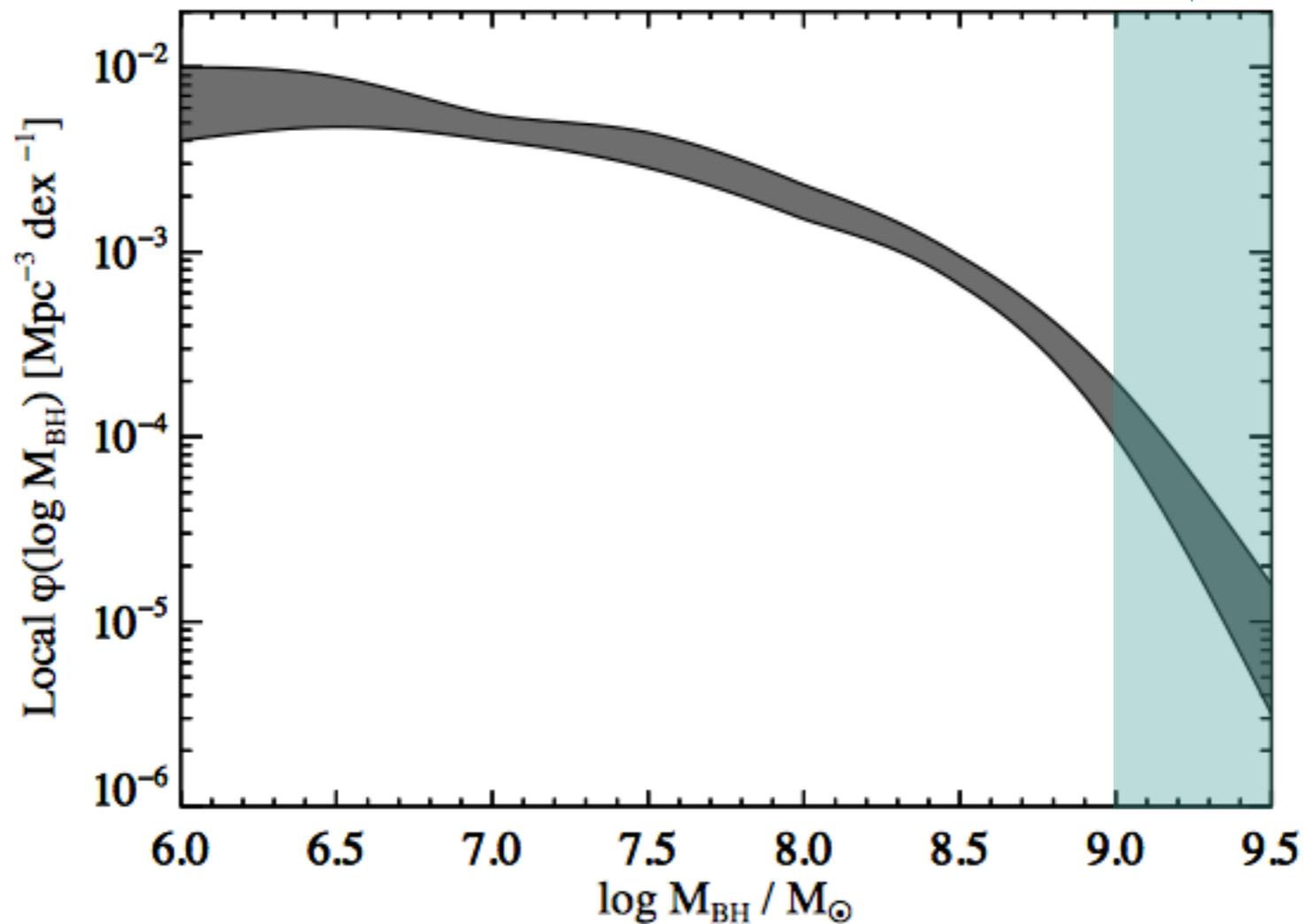
Fan+ 2000, 2001

Willott, McLure & Jarvis 2003

Mortlock+ 2011

Wu+ 2015

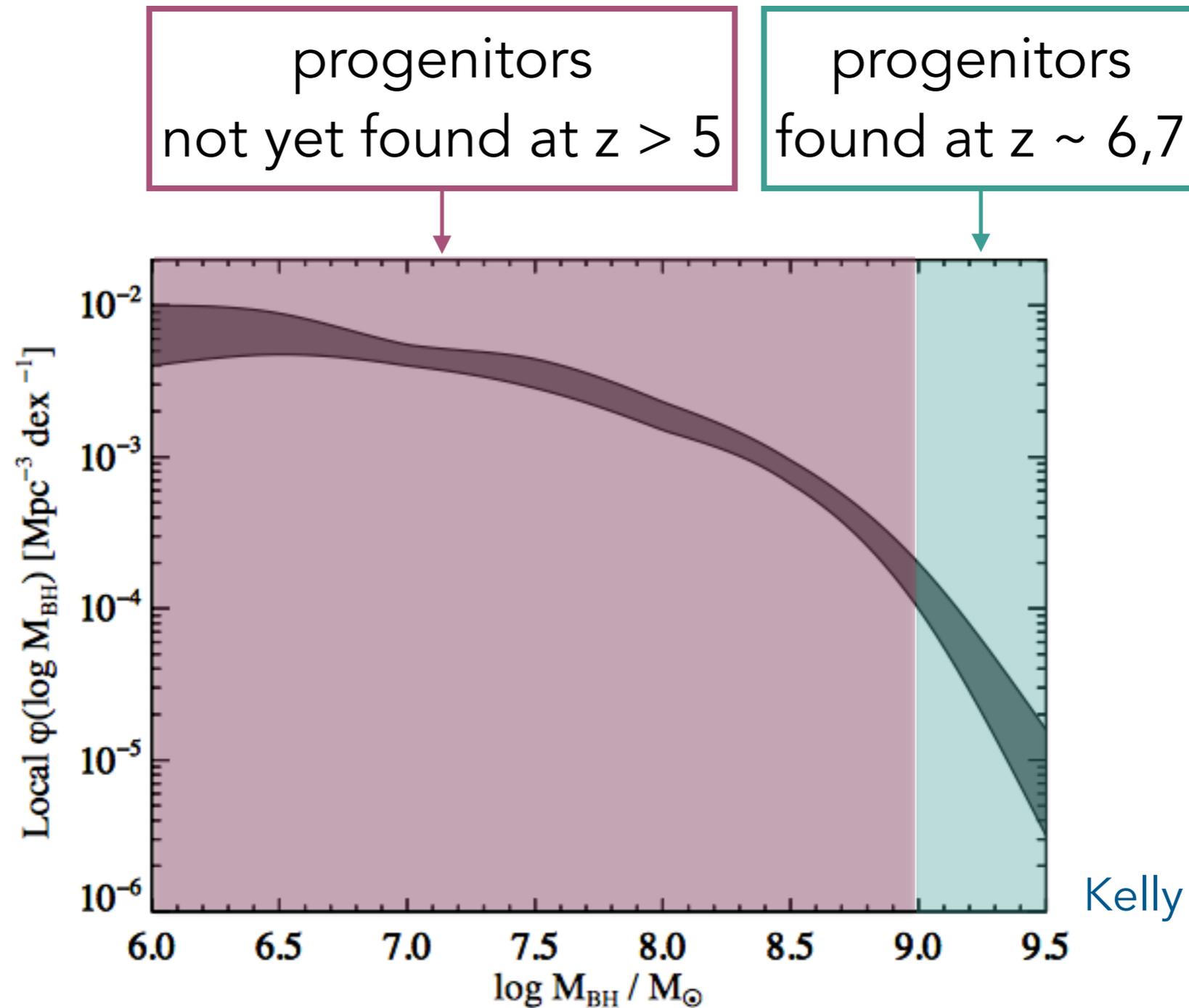
progenitors
found at $z \sim 6,7$



Kelly & Merloni, 2011

Motivation

Local BH mass function:



Kelly & Merloni, 2011

Expectations

▶ ~ 700 $z \sim 5$

Lyman Break Galaxies

[Stark+ 2009](#), [Vanzella+ 2009](#), [Wilkins+ 2010](#),
[Bouwens+ 2014](#), [Duncan+ 2014](#)

▶ at $z \sim 3$: AGN fraction $\sim 3\%$

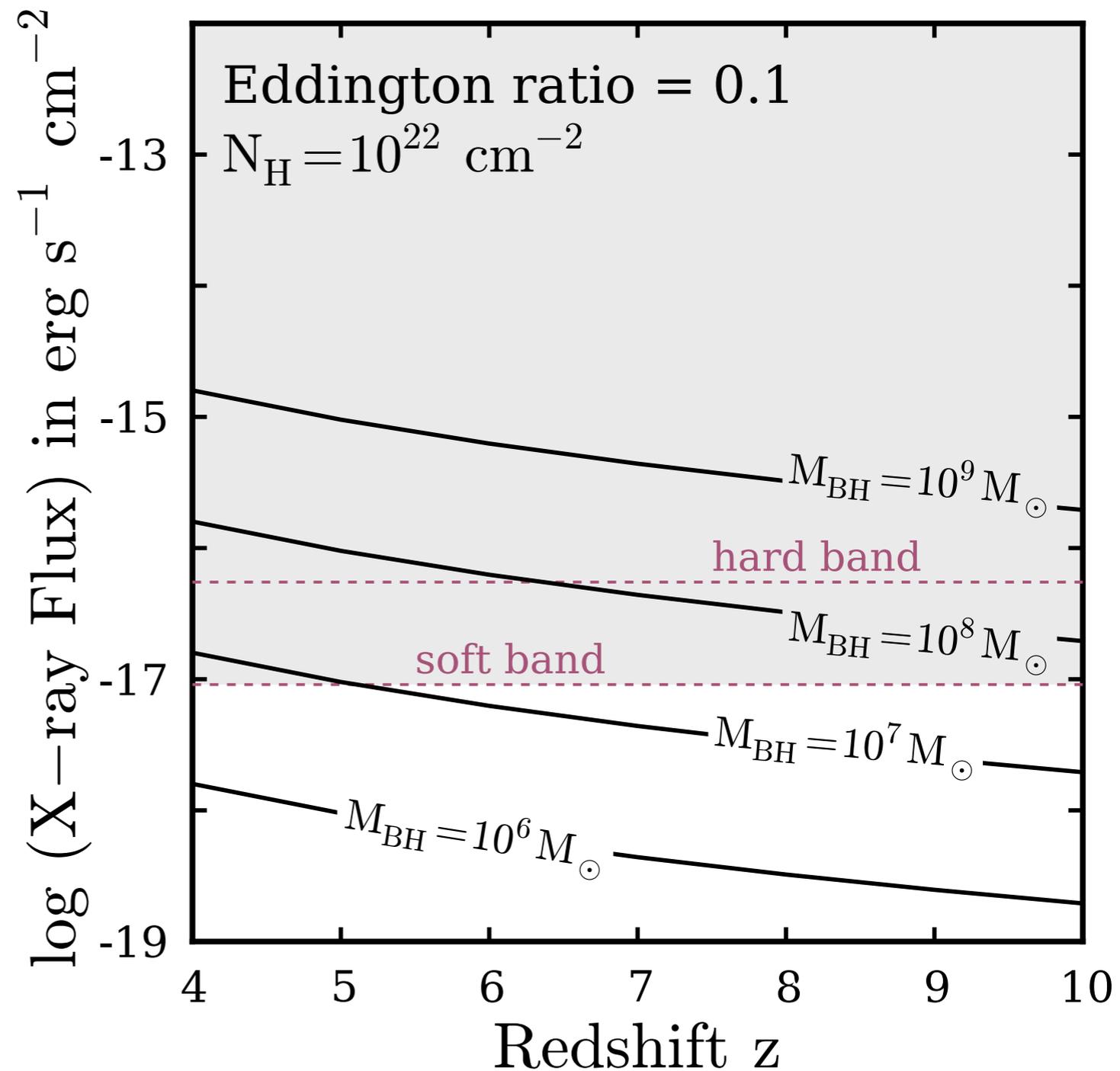
[Nandra+ 2002](#)

▶ expect to find ~ 20 $z \sim 5$

AGN

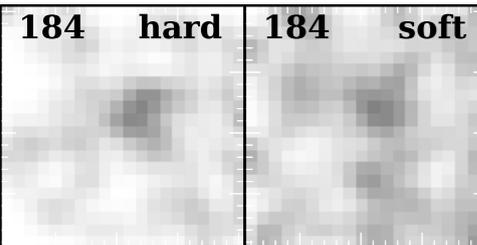
Expectations

- ▶ ~ 700 $z \sim 5$ Lyman Break Galaxies
- ▶ at $z \sim 3$: AGN fraction $\sim 3\%$
- ▶ expect to find ~ 20 $z \sim 5$ AGN



Data

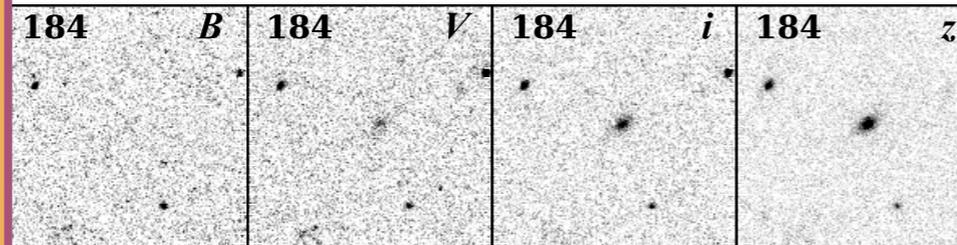
X-rays



Chandra/ACIS,
4Ms-catalog

Xue+ 2011

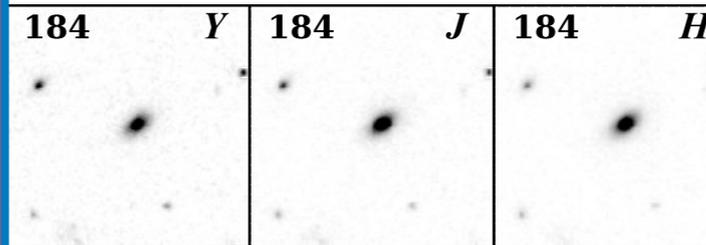
optical



HST/ACS, GOODS

Giavalisco+ 2004

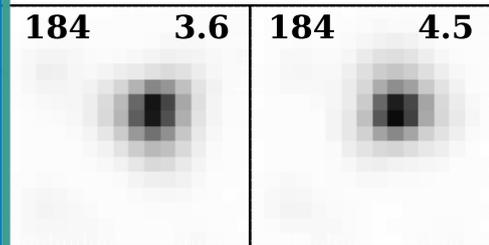
NIR



HST/WFC3, CANDELS

Grogin+ 2011,
Koekemoer+ 2011

IR



Spitzer/IRAC,
SIMPLE

van Dokkum+ 2005,
Damen+ 2011

Redshift tests

visual class.

stacking

colour
criteria

photo-z

Hardness
Ratio

Redshift tests

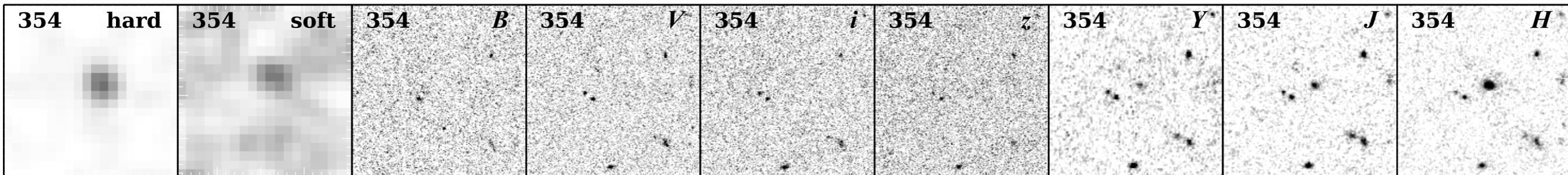
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Steidel+ 1999

Giavalisco 2002

Dunlop 2013

Redshift tests

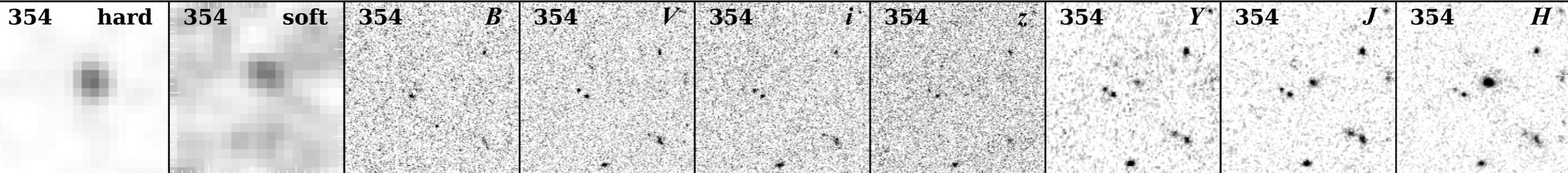
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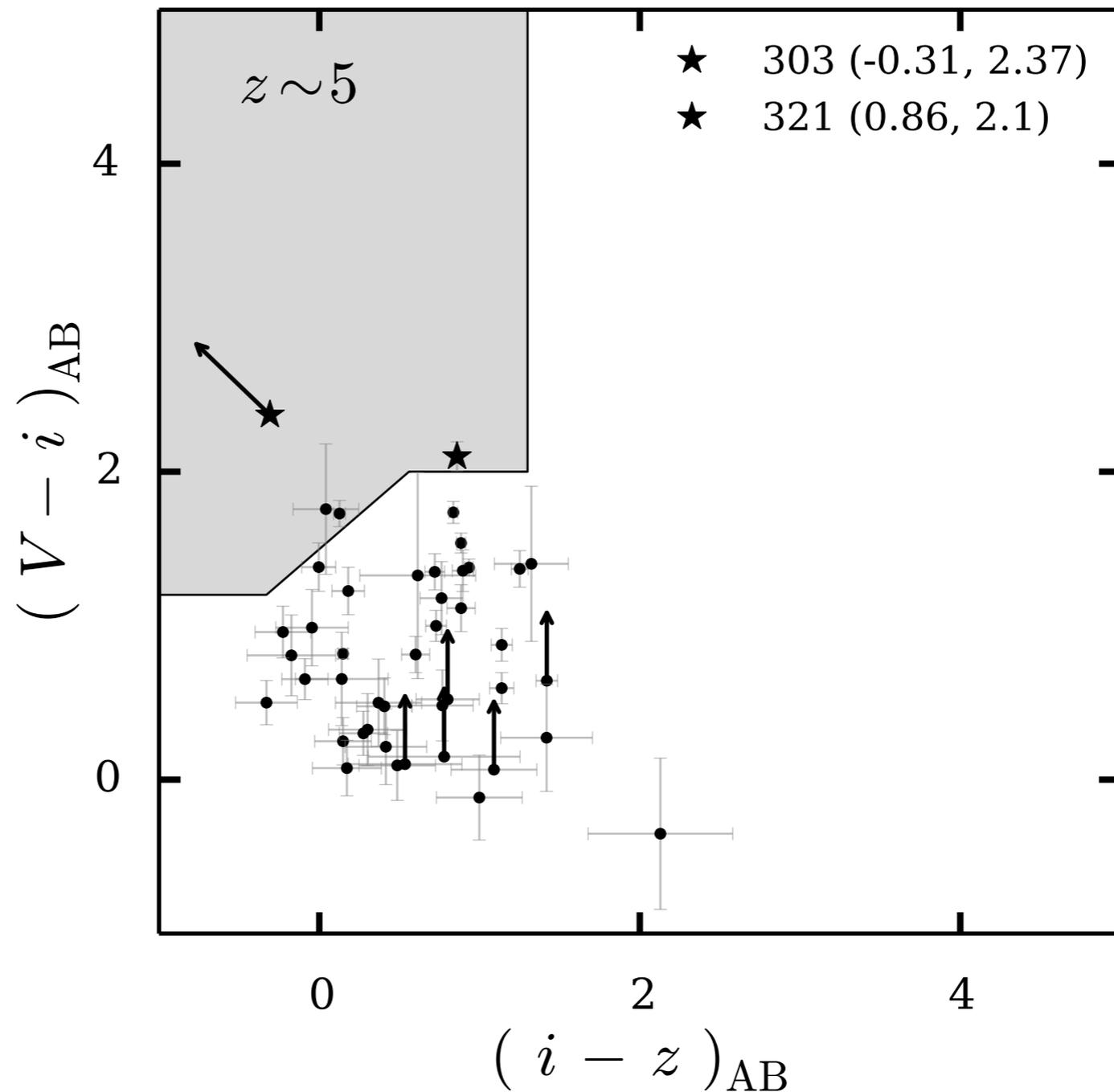
Vanzella+ 2009

Steidel+ 1999

Daddi+ 2004

van Dokkum+ 2003

Taniguchi+ 2005



Redshift tests

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stacking

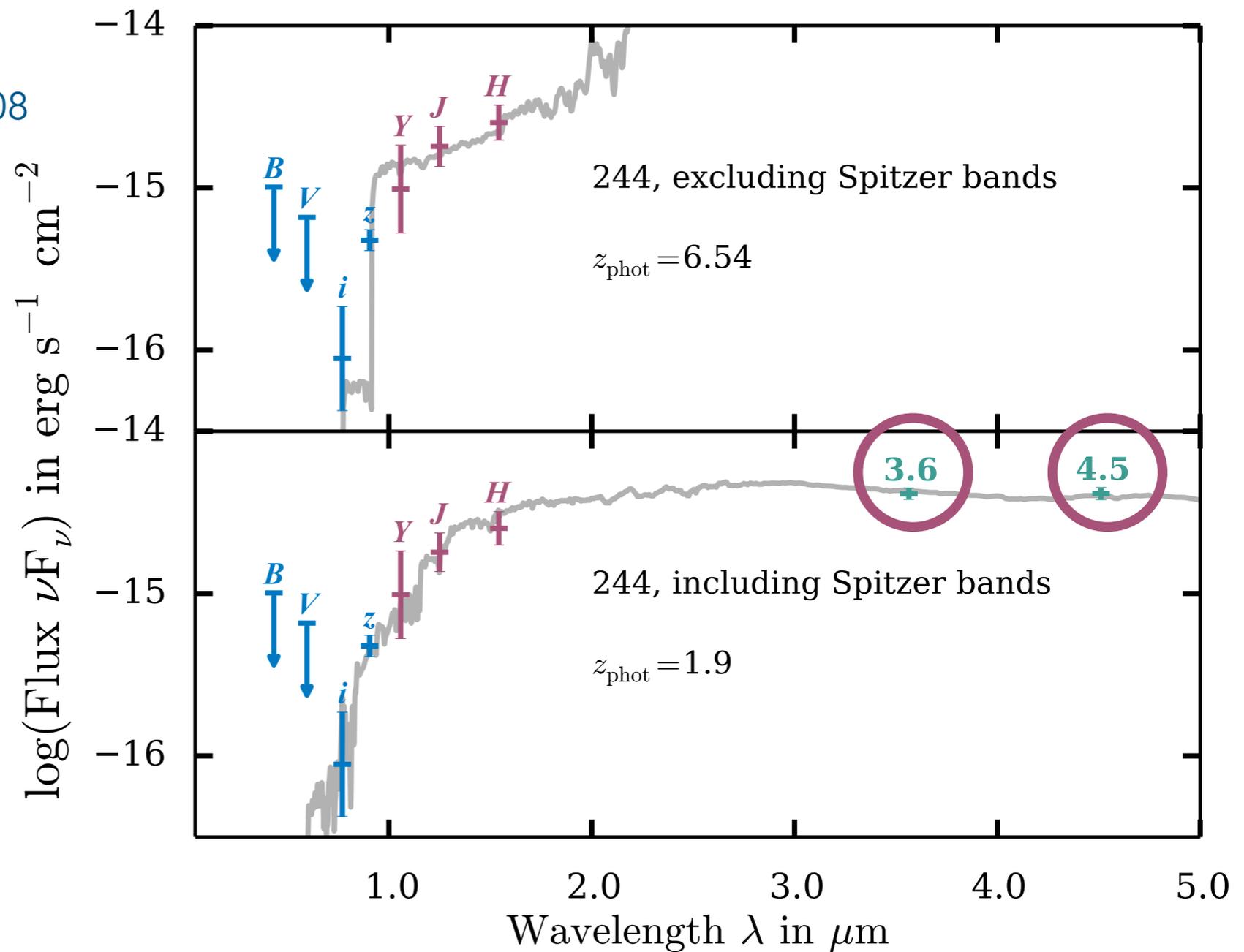
colour
criteria

photo-z

Hardness
Ratio

EAZY:

Brammer, van Dokkum & Coppi 2008



Redshift tests

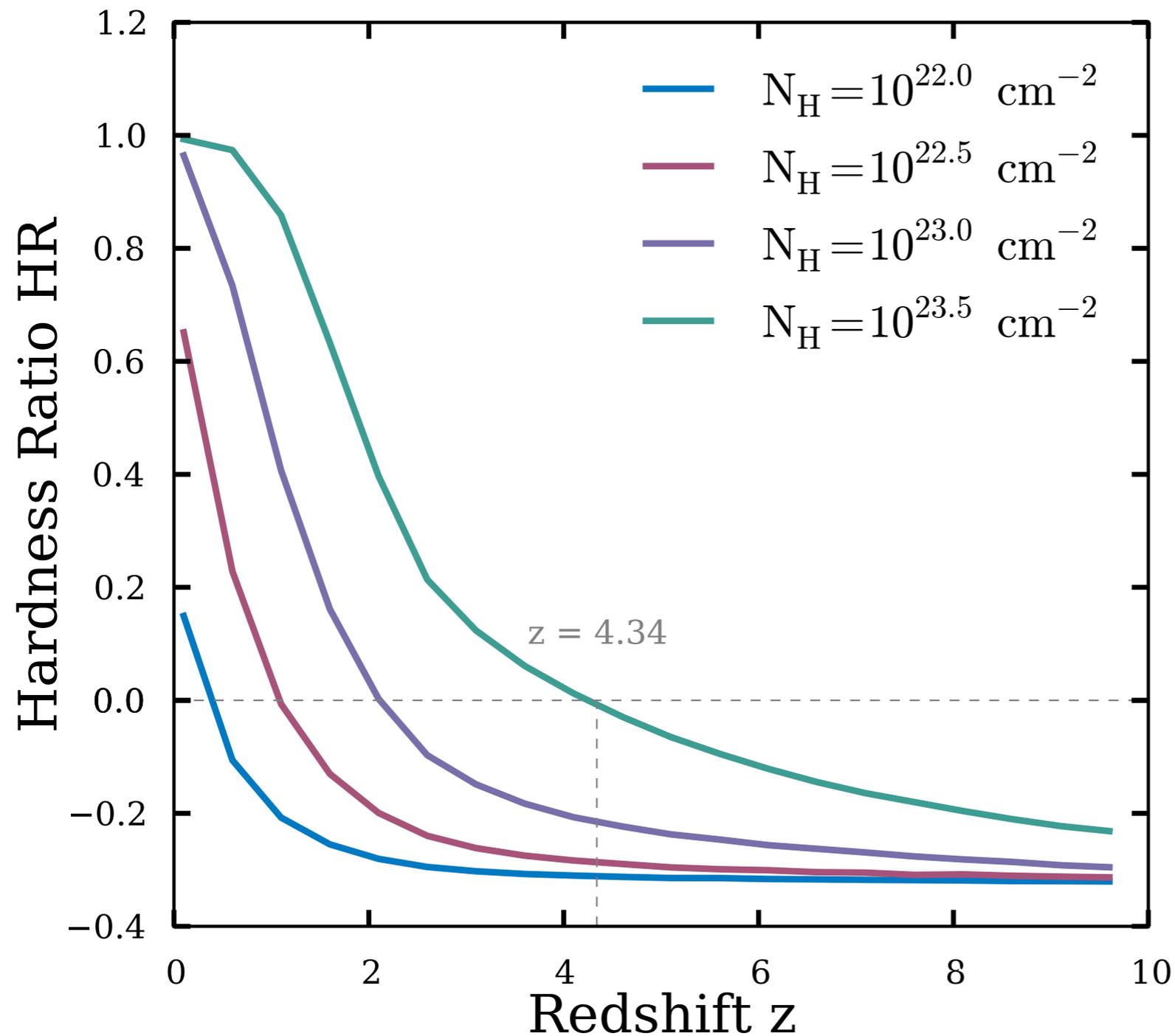
visual class.

stacking

colour
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photo-z

Hardness
Ratio



$$HR = \frac{H - S}{H + S}$$

Wang+ 2004
Arnaud+ 1996

Redshift tests

visual class.

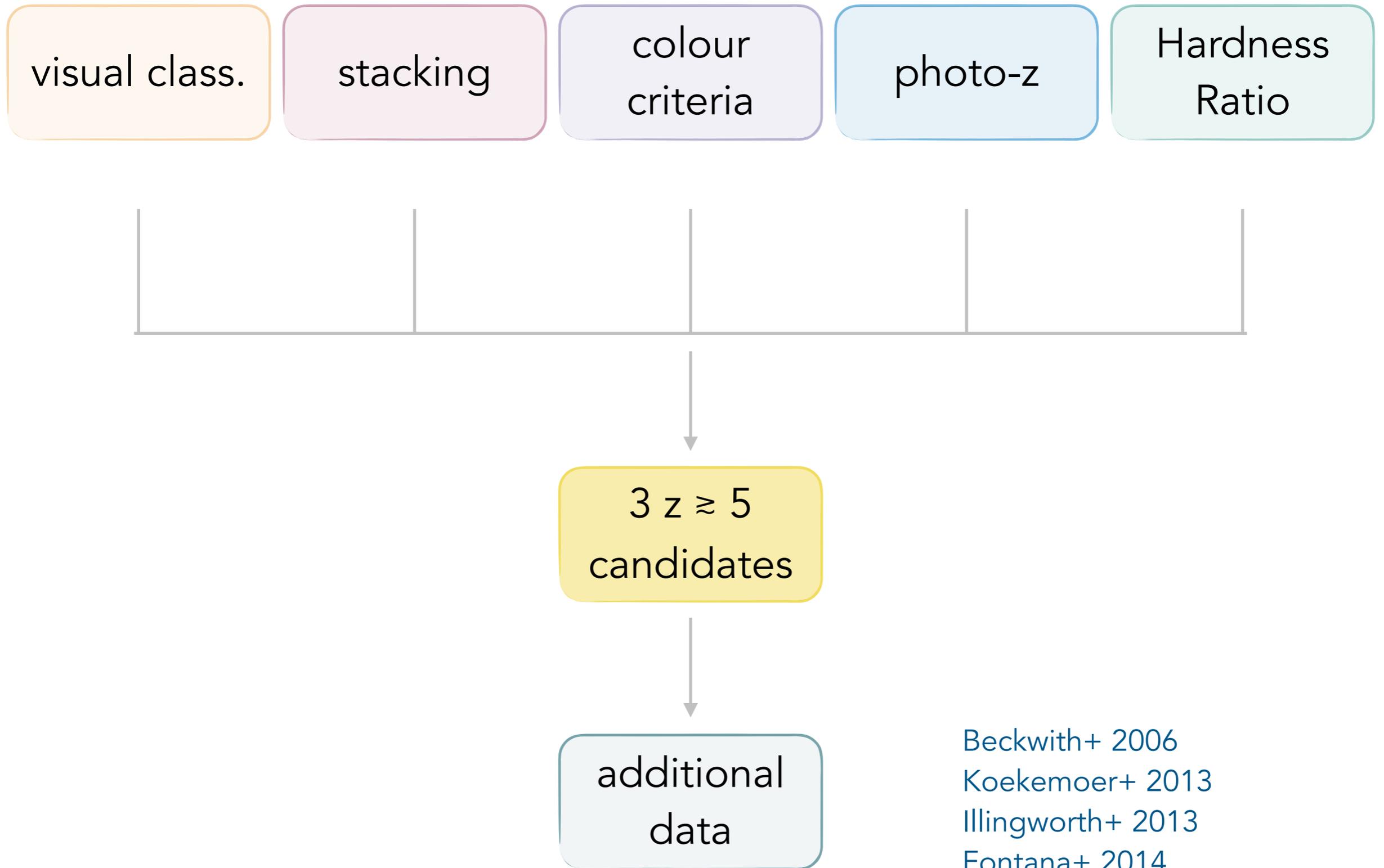
stacking

colour
criteria

photo-z

Hardness
Ratio

Redshift tests



Beckwith+ 2006
Koekemoer+ 2013
Illingworth+ 2013
Fontana+ 2014

No convincing $z \gtrsim 5$ AGN candidates
in the *Chandra* Deep Field South

Possible explanations

physical reasons

- ▶ BH occupation fraction
- ▶ AGN fraction
- ▶ super-Eddington

$$N_{\text{AGN}} = f_{\text{active}} \times f_{\text{occ}} \times N_{\text{galaxies}}$$

Volonteri & Silk 2014, Madau, Haardt & Dotti 2014,
Alexander & Natarajan 2014

Possible explanations

physical reasons

- ▶ BH occupation fraction
- ▶ AGN fraction
- ▶ super-Eddington

observational reasons

- ▶ optically faint galaxies

Possible explanations

physical reasons

- ▶ BH occupation fraction
- ▶ AGN fraction
- ▶ super-Eddington

observational reasons

- ▶ optically faint galaxies
- ▶ X-ray obscuration

Possible explanations

physical reasons

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observational reasons

- ▶ optically faint galaxies
- ▶ X-ray obscuration
- ▶ BH - BH coalescence

Possible explanations

physical reasons

- ▶ BH occupation fraction
- ▶ AGN fraction
- ▶ super-Eddington

*2.8 Ms Chandra COSMOS
Legacy Survey,
Francesca Civano*

observational reasons

- ▶ optically faint galaxies
- ▶ X-ray obscuration
- ▶ BH - BH coalescence

7 Ms Chandra survey, Niel Brandt

Summary

- ▶ searched for $z \gtrsim 5$ AGN in the *Chandra* Deep Field South
- ▶ *Chandra* 4-MS + GOODS, CANDELS, *Spitzer*
- ▶ 5 redshift tests
- ▶ no convincing $z \gtrsim 5$ AGN candidates
- ▶ explanations: low BH occupation/AGN fraction, super Eddington growth, optically faint/obscured galaxies, BH - BH coalescence