

Clustering Measurements of $M_{\text{halo}}(M_{\star})$ in CANDELS

Catherine White

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Advisor: Harry Ferguson





Halo bias $b_h(M_{\text{halo}})$

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+

Halo occupation model

Halo bias $b_h(M_{\text{halo}})$

+

Halo occupation model

+

Observed galaxy bias

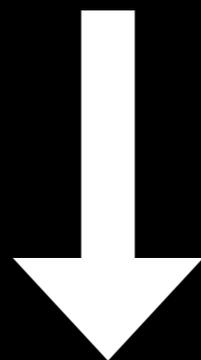
Halo bias $b_h(M_{\text{halo}})$

+

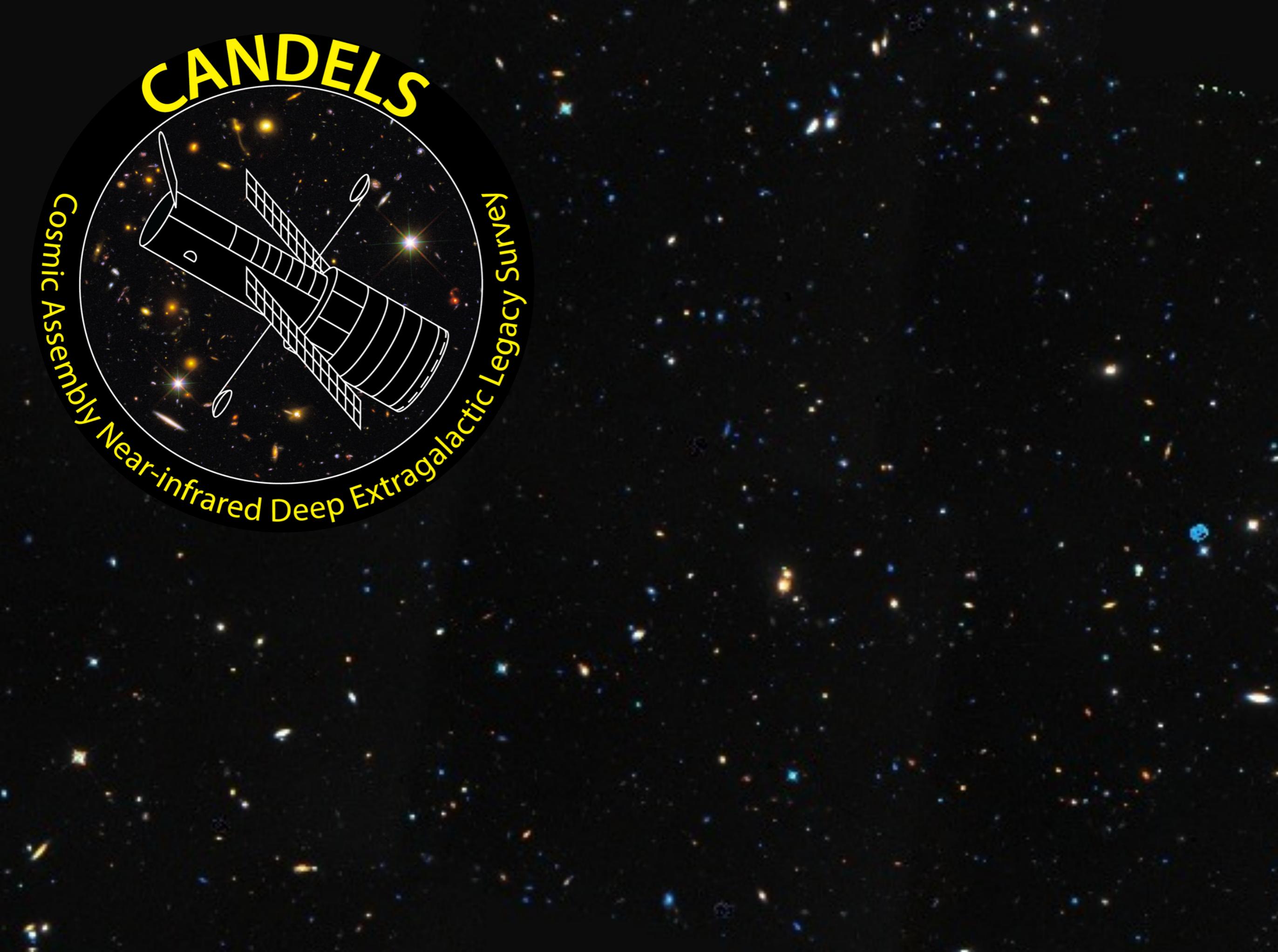
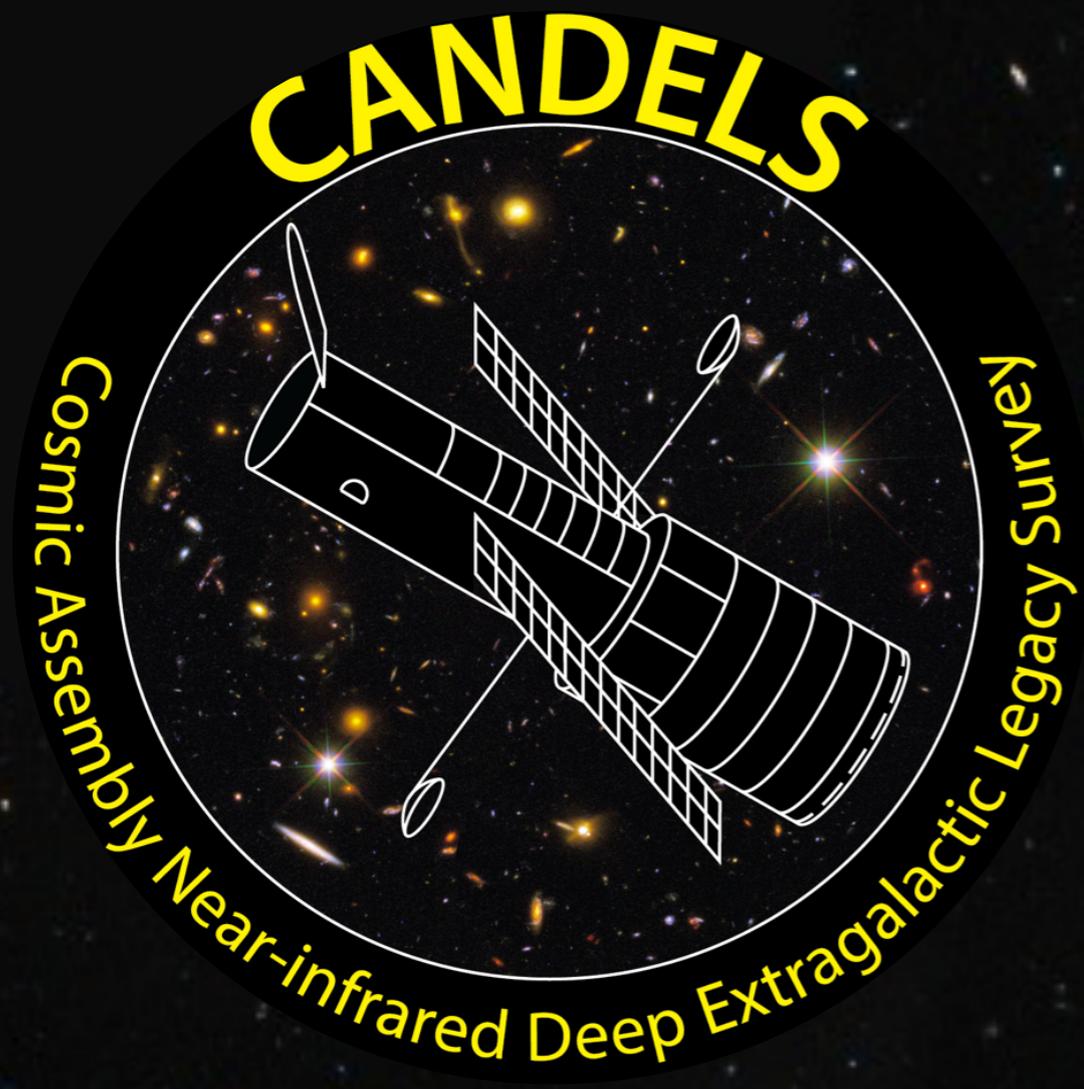
Halo occupation model

+

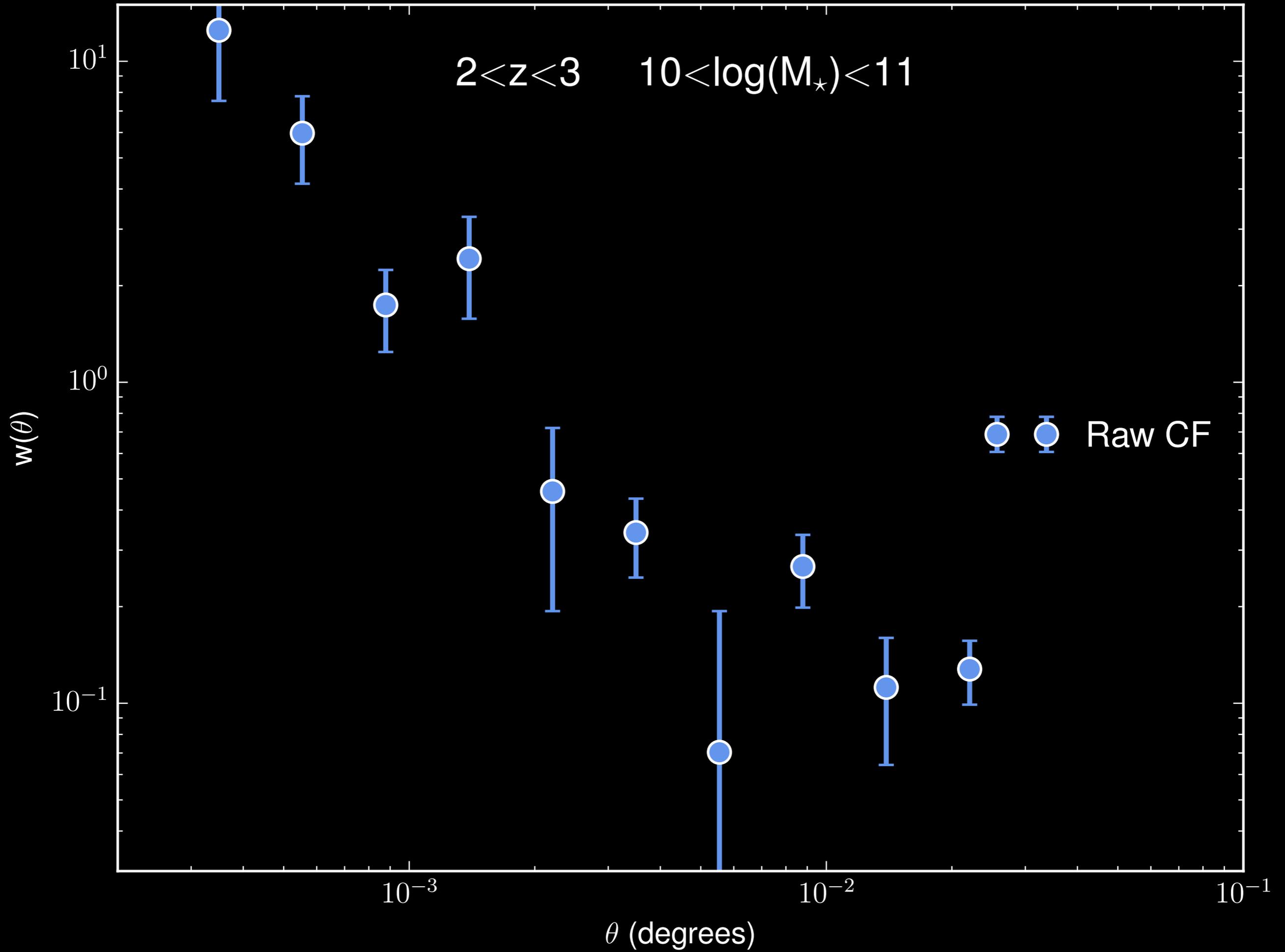
Observed galaxy bias

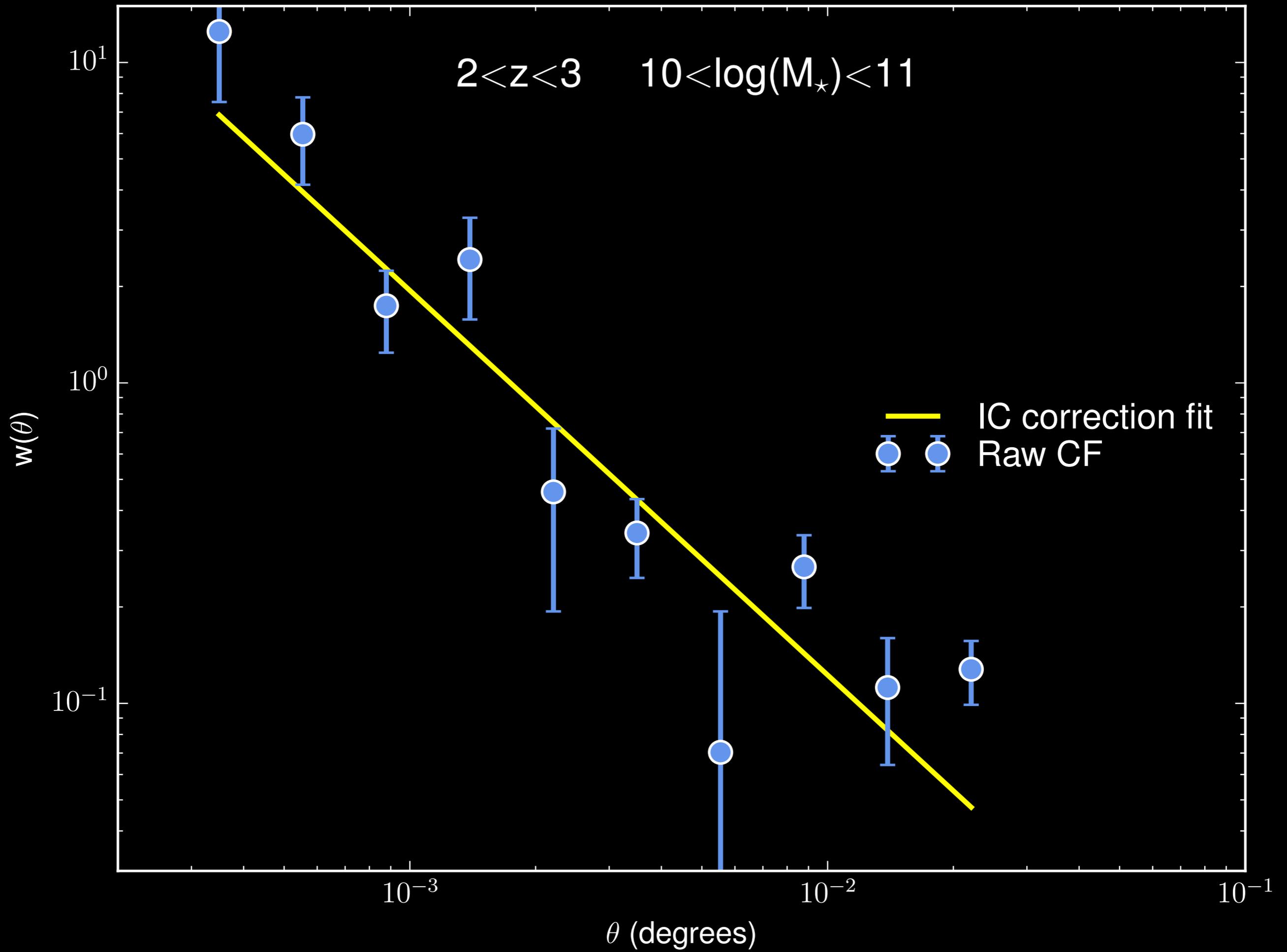


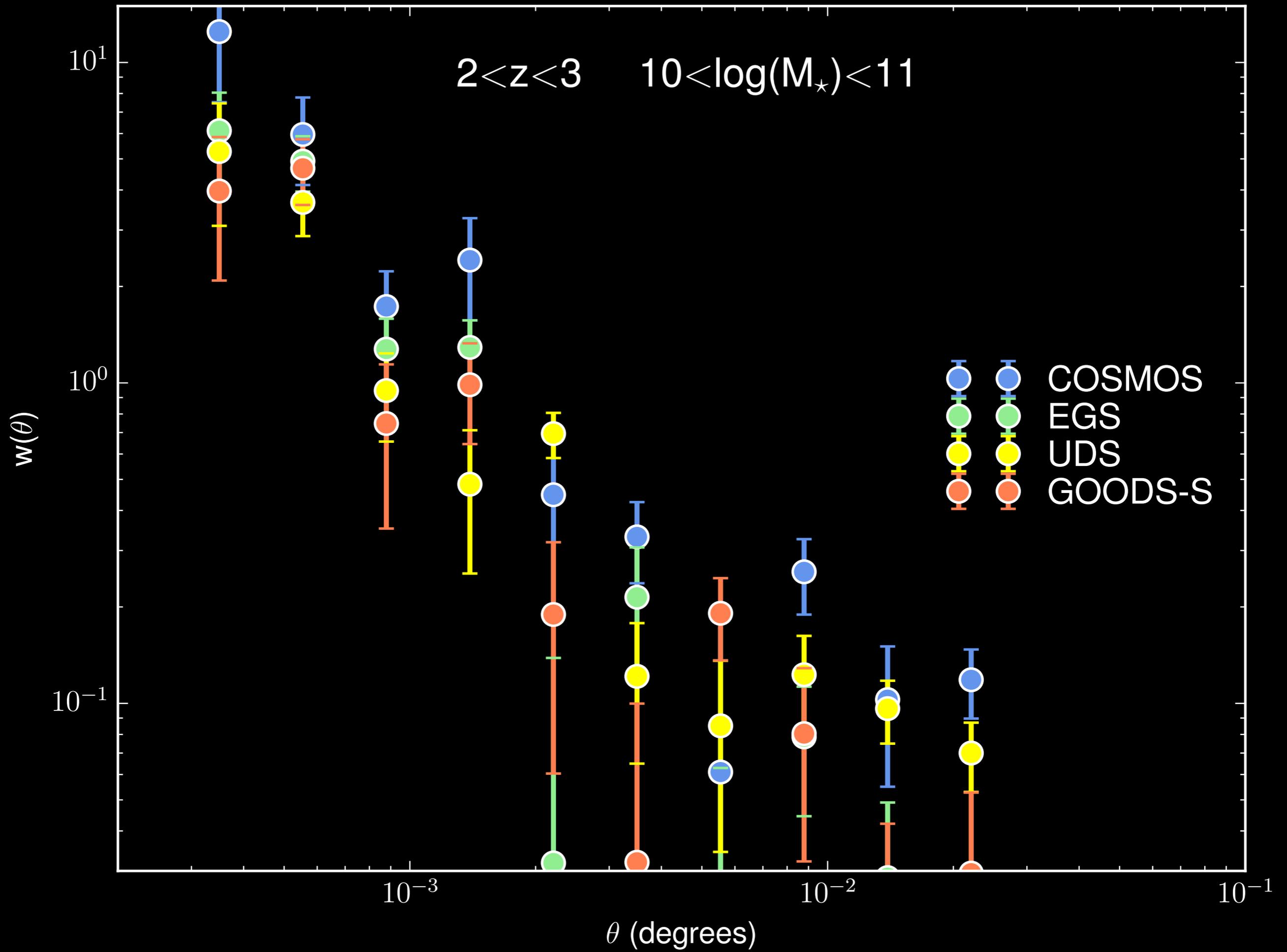
Halo mass

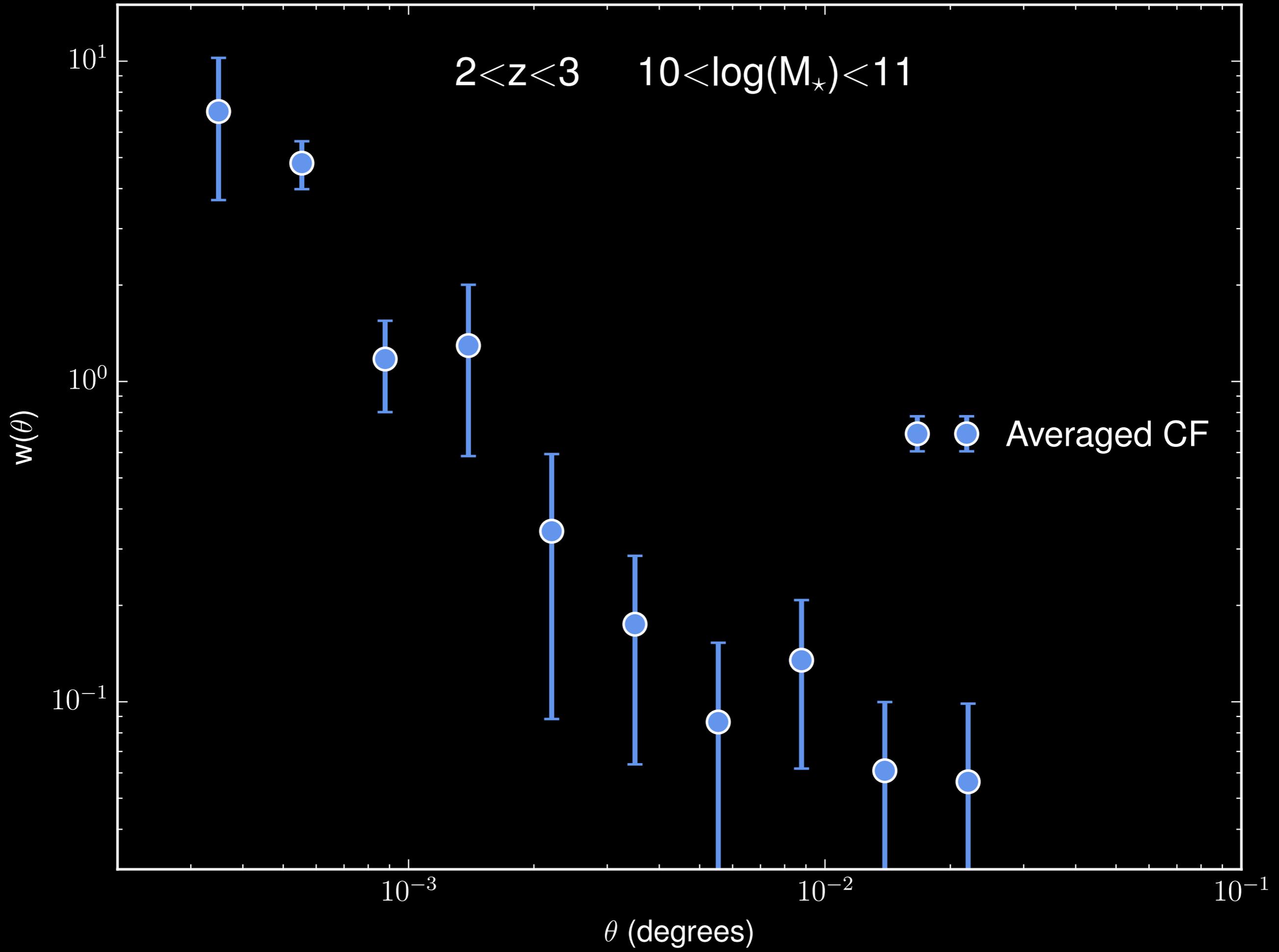


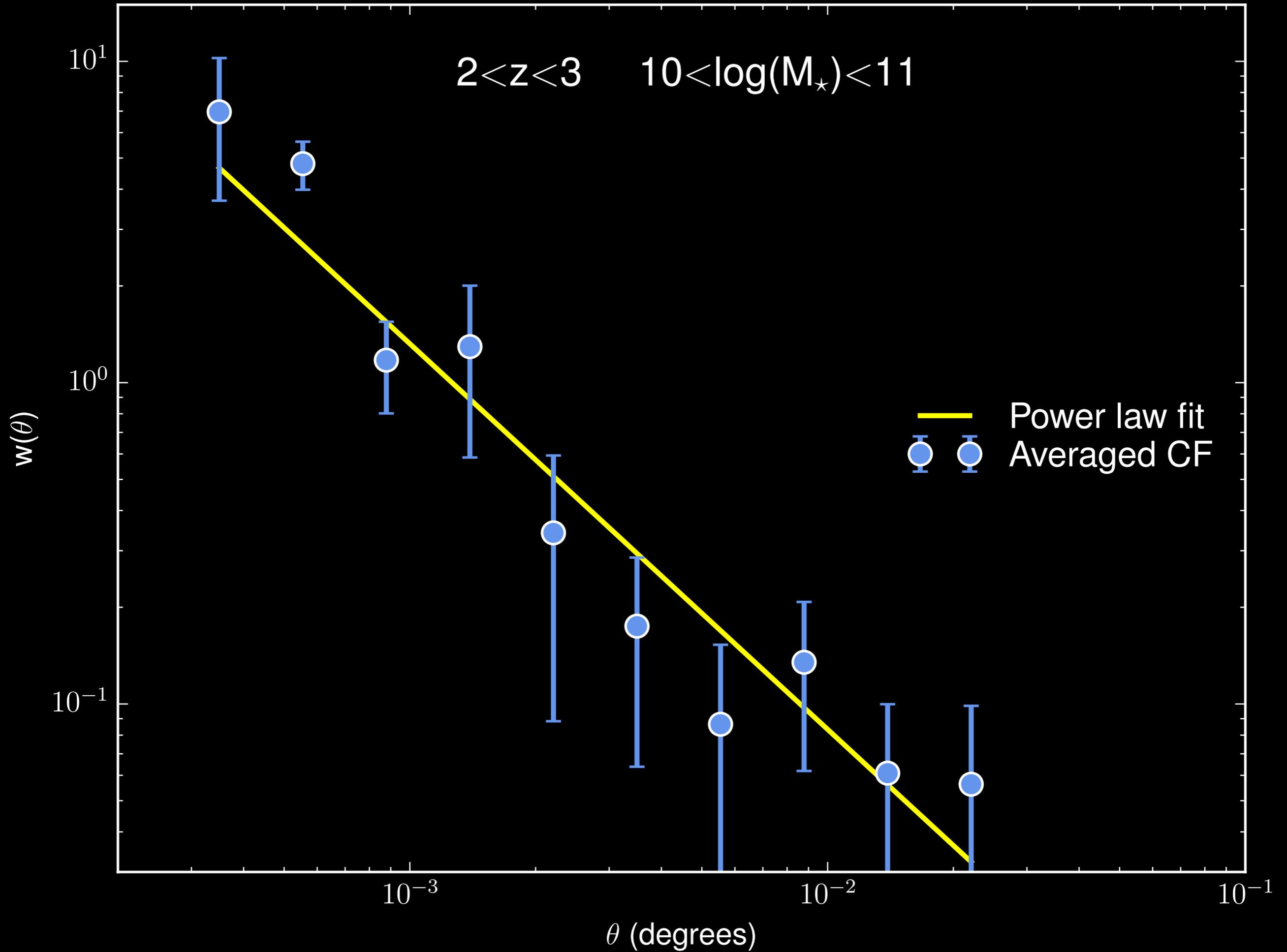
Process





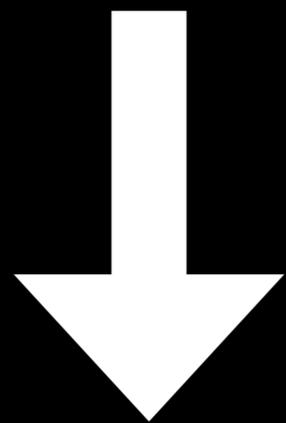






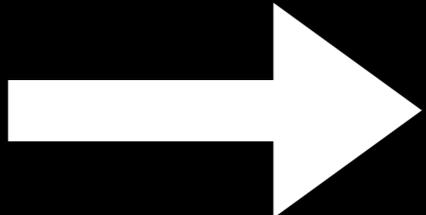
$$w(\theta) = A \theta^{-\beta}$$

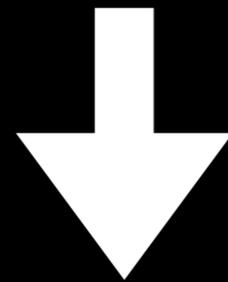
$$w(\theta) = A \theta^{-\beta}$$



$$\xi(r) = (r/r_0)^{-\gamma}$$

$r_0, \gamma \rightarrow \text{bias}$

r_0, γ  bias



halo mass

Uncertainties

Integral constraint

Uncertainties

Integral constraint

Power law fit

Uncertainties

Integral constraint

Power law fit

$N(z)$ for Limber equation

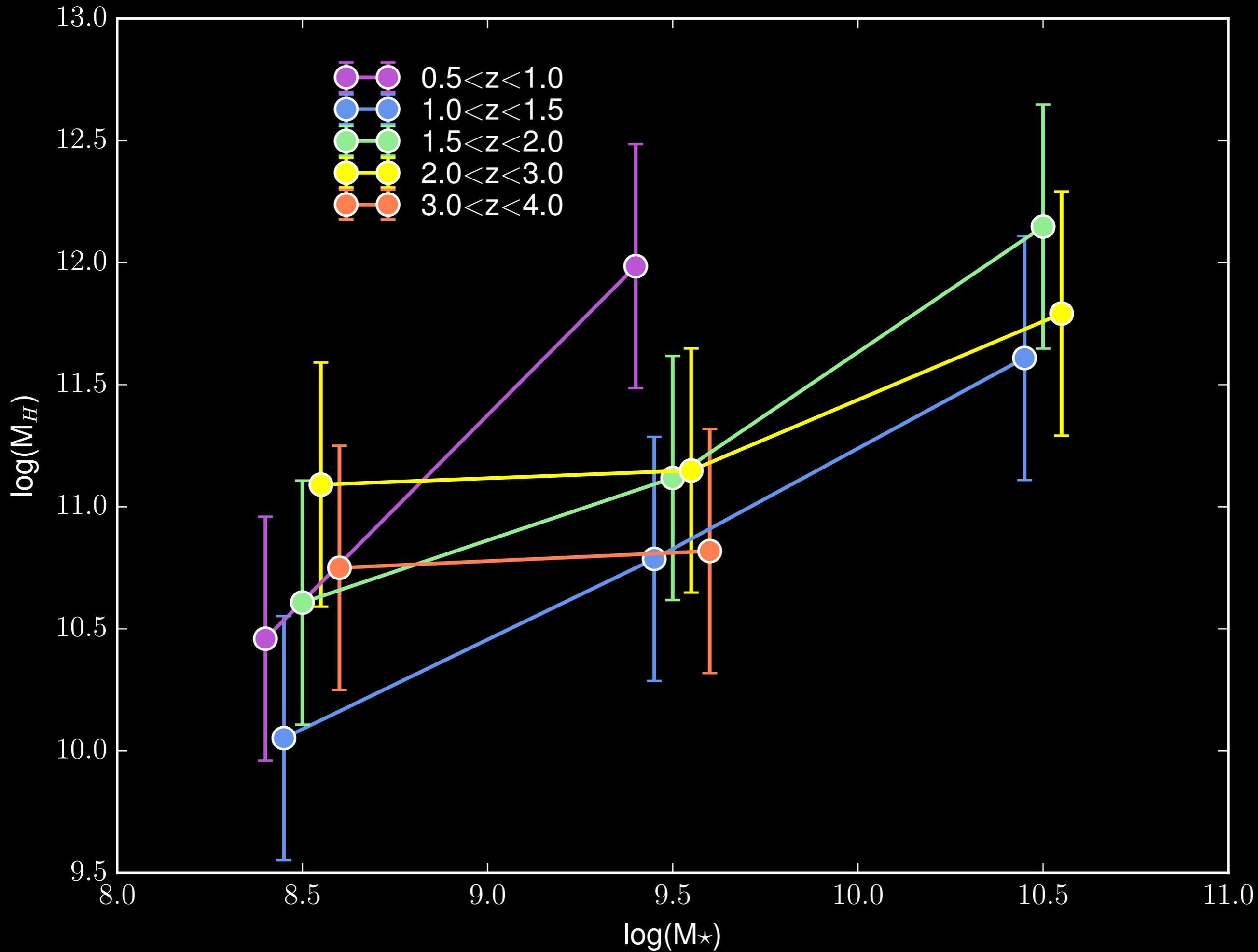
Uncertainties

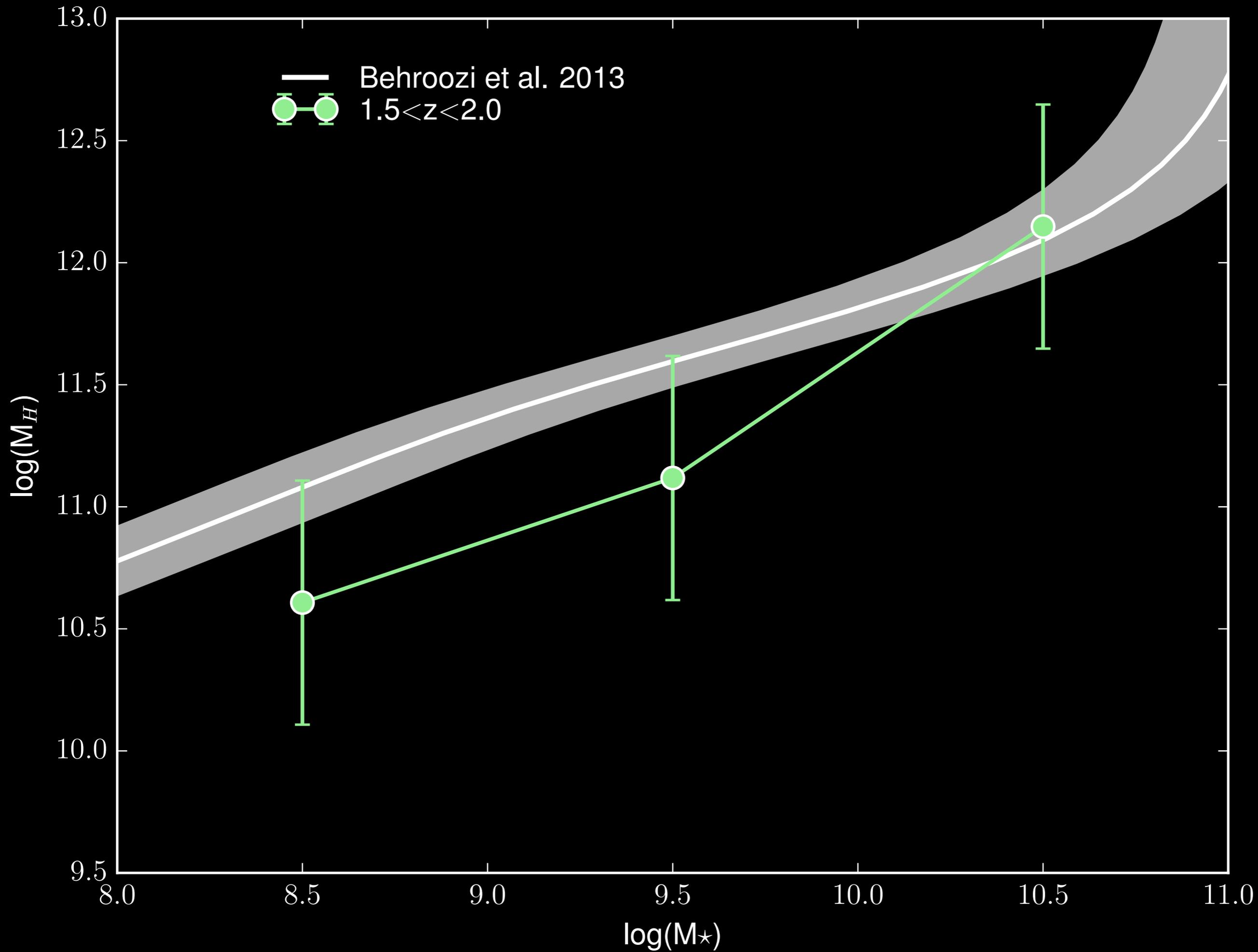
Integral constraint

Power law fit

$N(z)$ for Limber equation

Bias to halo mass





Going Forward

Improve integral constraint / fitting

Going Forward

Improve integral constraint / fitting

Proper error computation

Going Forward

Improve integral constraint / fitting

Proper error computation

Use halo occupation distribution

Going Forward

Improve integral constraint / fitting

Proper error computation

Use halo occupation distribution

More science!

Collaborators

Harry Ferguson, Jeff Newman, Peter Behroozi,
Casey Papovich, Steve Finkelstein, Christina
Williams, Guillermo Barro, Rachel Somerville

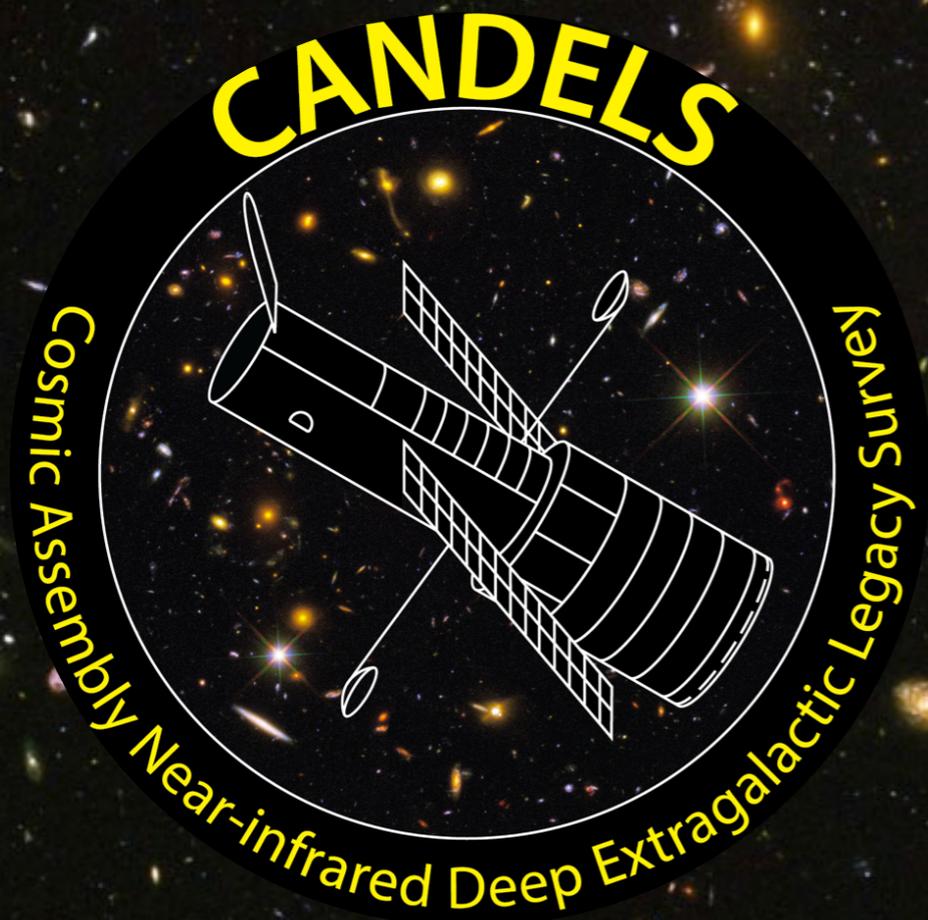


Image Sources

<http://hubblesite.org/>

<http://candels.ucolick.org/>

