

# THE EXTRA-GALACTIC UNIVERSE AND ITS COSMOLOGY

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PORTUGUESE ALMA COMMUNITY DAY 2015

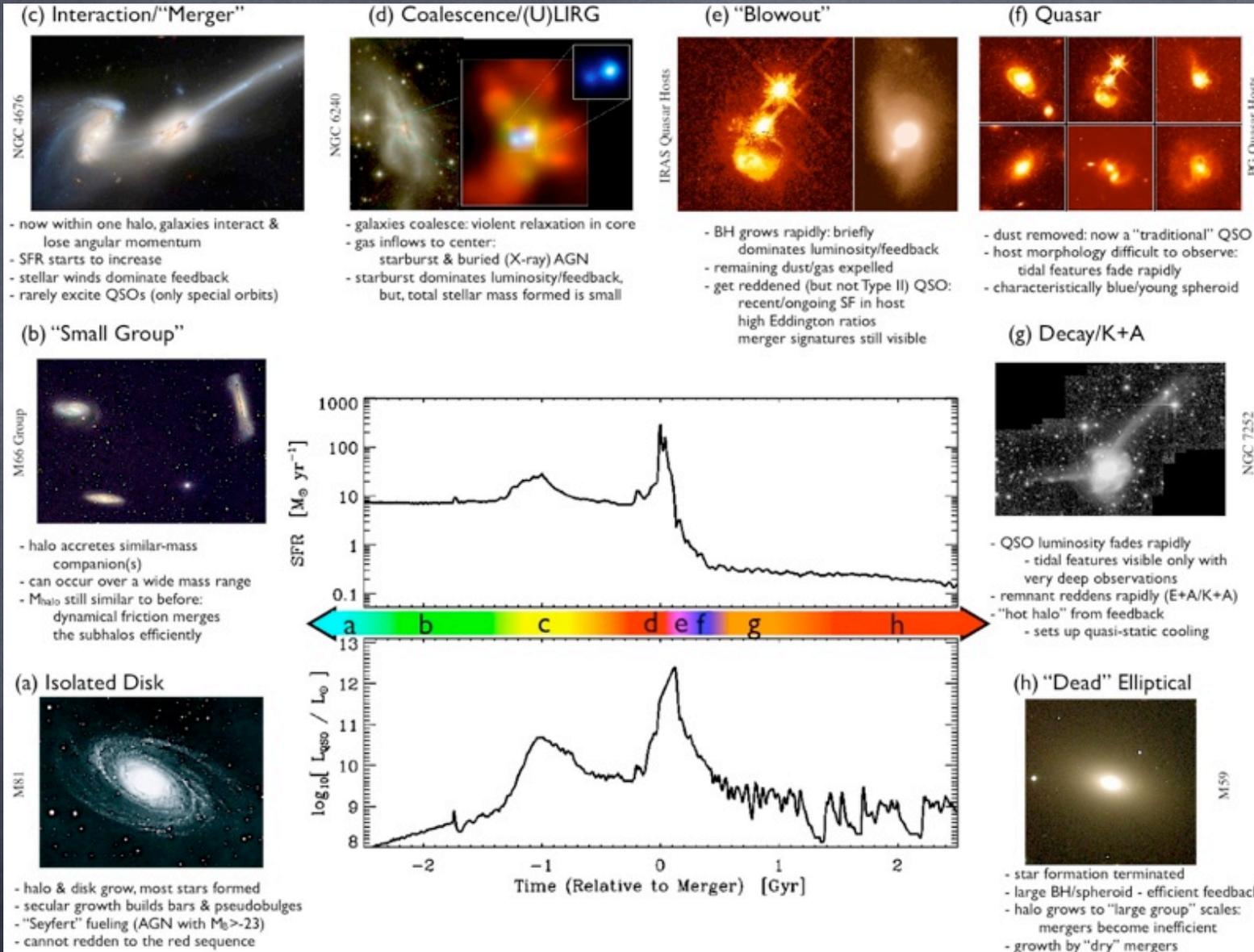
IA-LISBOA, 24<sup>TH</sup> FEBRUARY 2015

**ia**



instituto de astrofísica  
e ciências do espaço

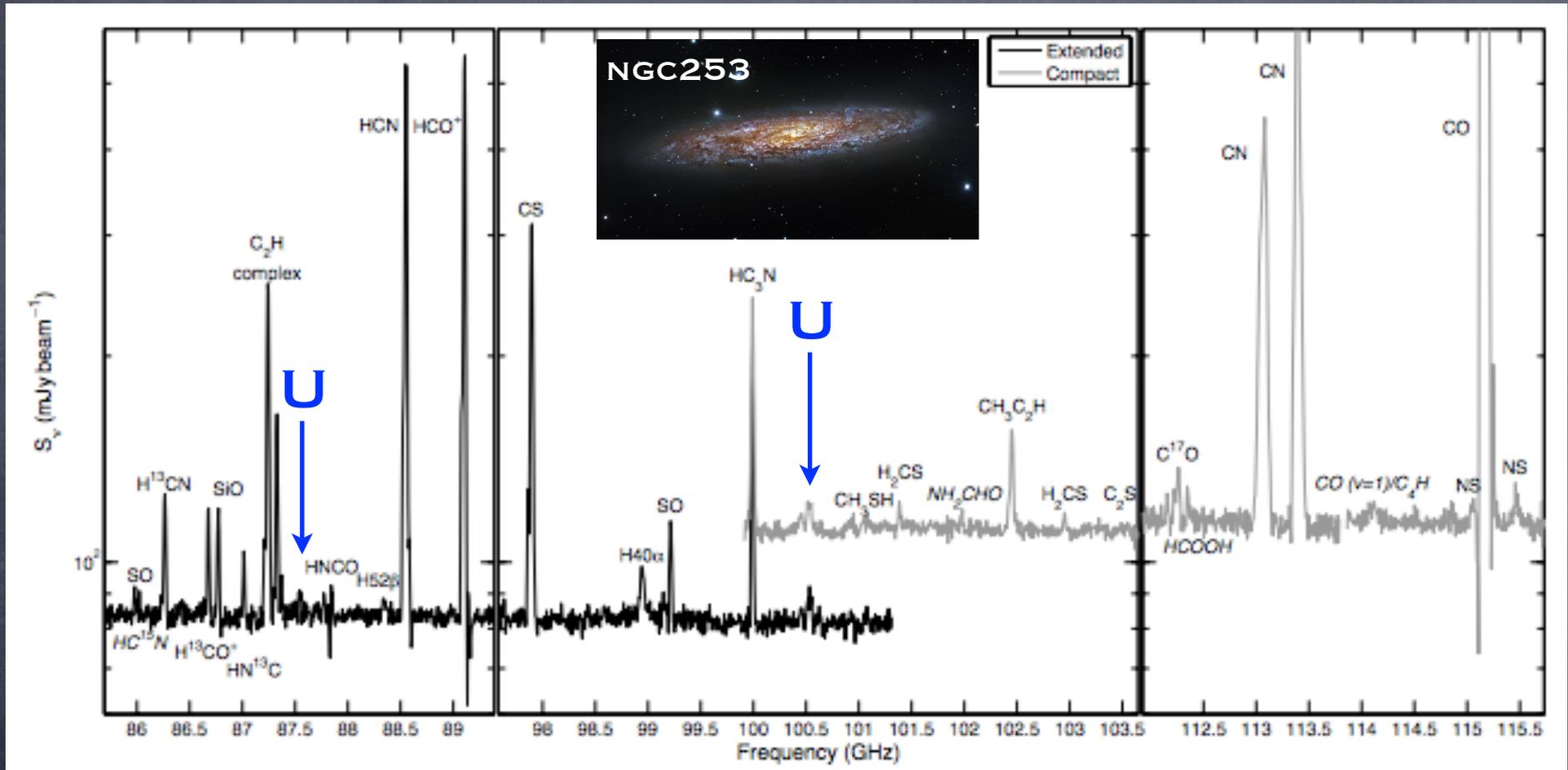




P. HOPKINS ET AL.(2008, FIG. 1)

alma can study  
 all stages

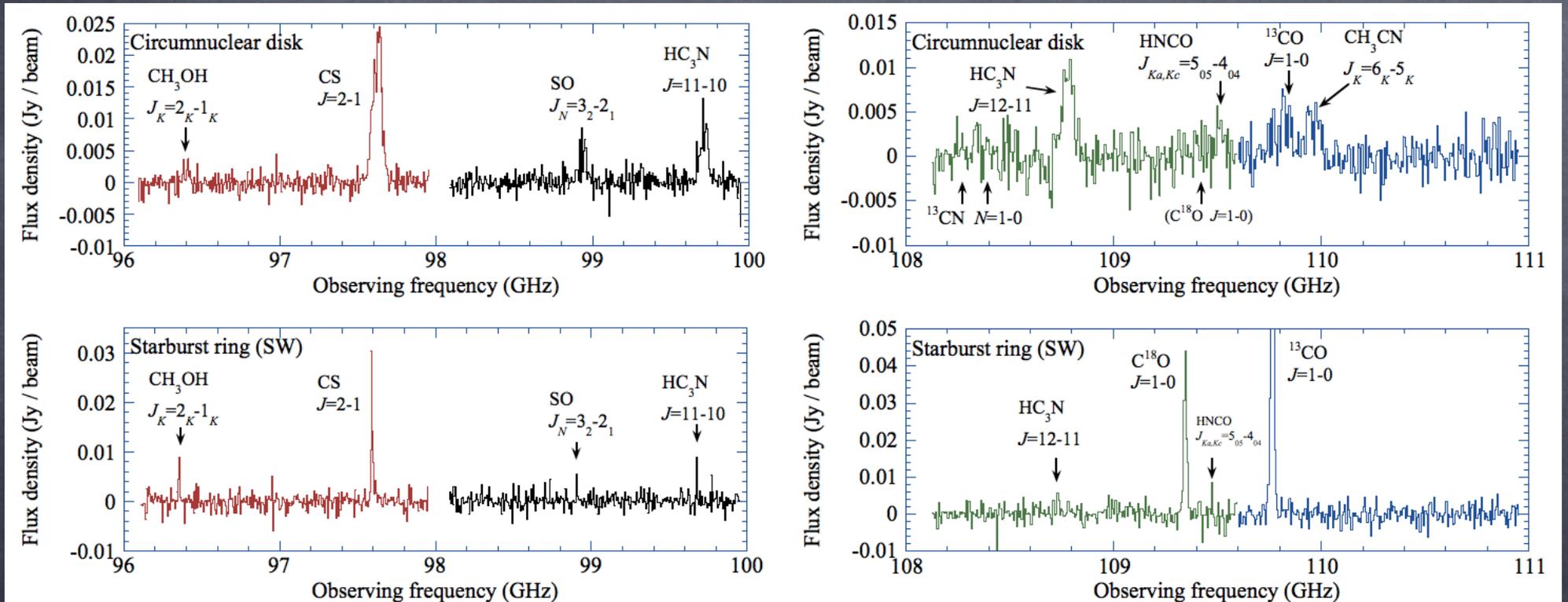
# LOCAL UNIVERSE



MEIER ET AL. (2015, FIG.2)

- spectral line detection
- unidentified species
- normalization real
- Cyc0

# LOCAL UNIVERSE

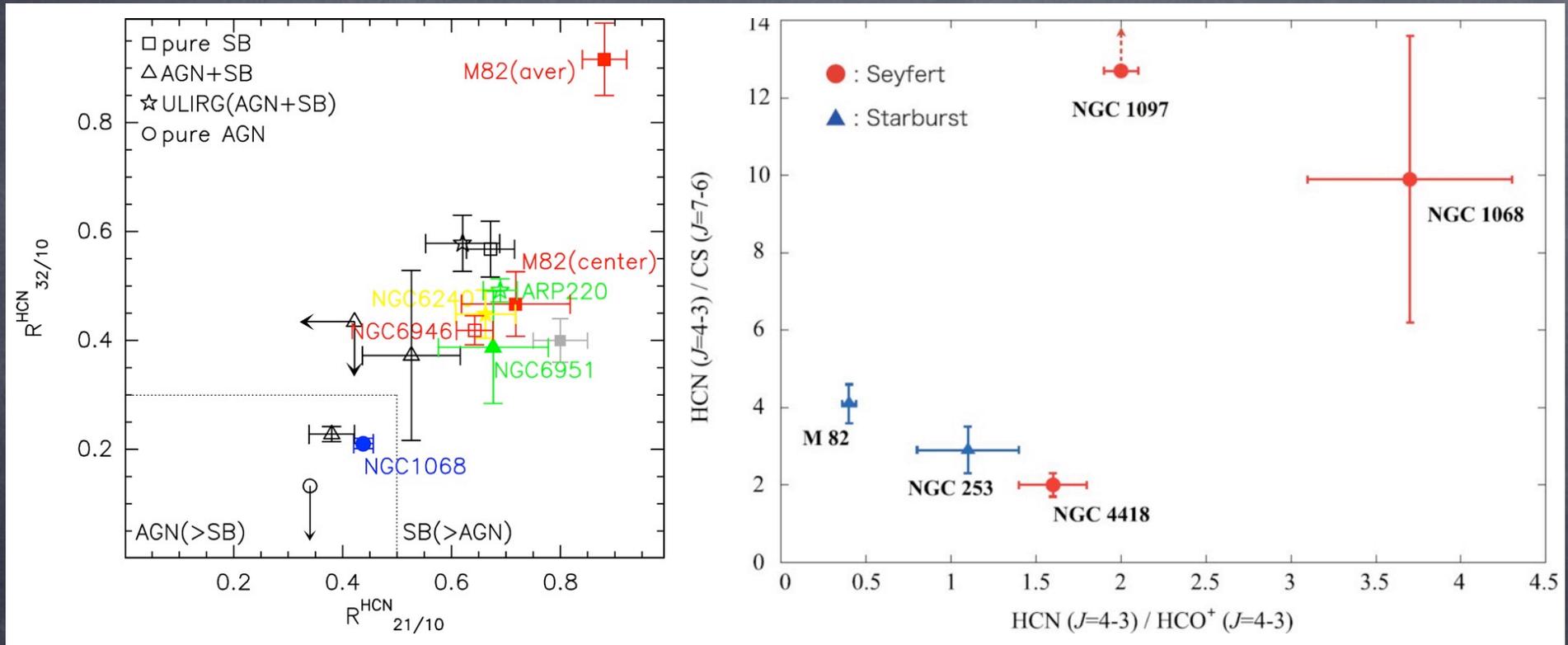


TAKANO ET AL. (2014, FIG.3)

spatially  
 resolved  
 spectroscopy  
 Cyc0



# LOCAL UNIVERSE

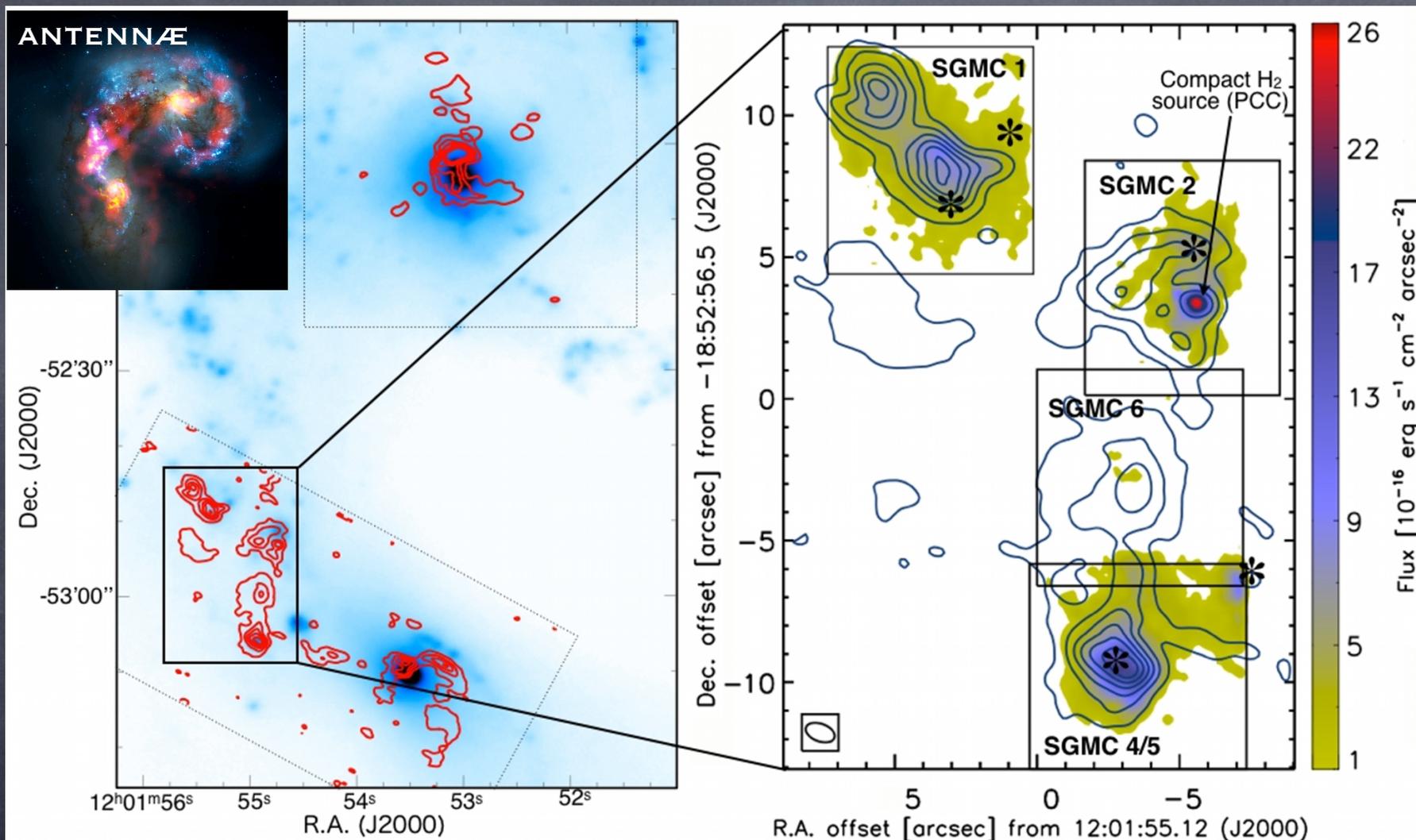


KRIPS ET AL. (2008, FIG. 3)  
 [SEE ALSO MEIJERINK ET AL. (2007)]

IZUMI ET AL. (2013, FIG. 11)

mm-bpt diagram  
 need for spatial  
 resolution though  
 Izumi, Cyc0, B7

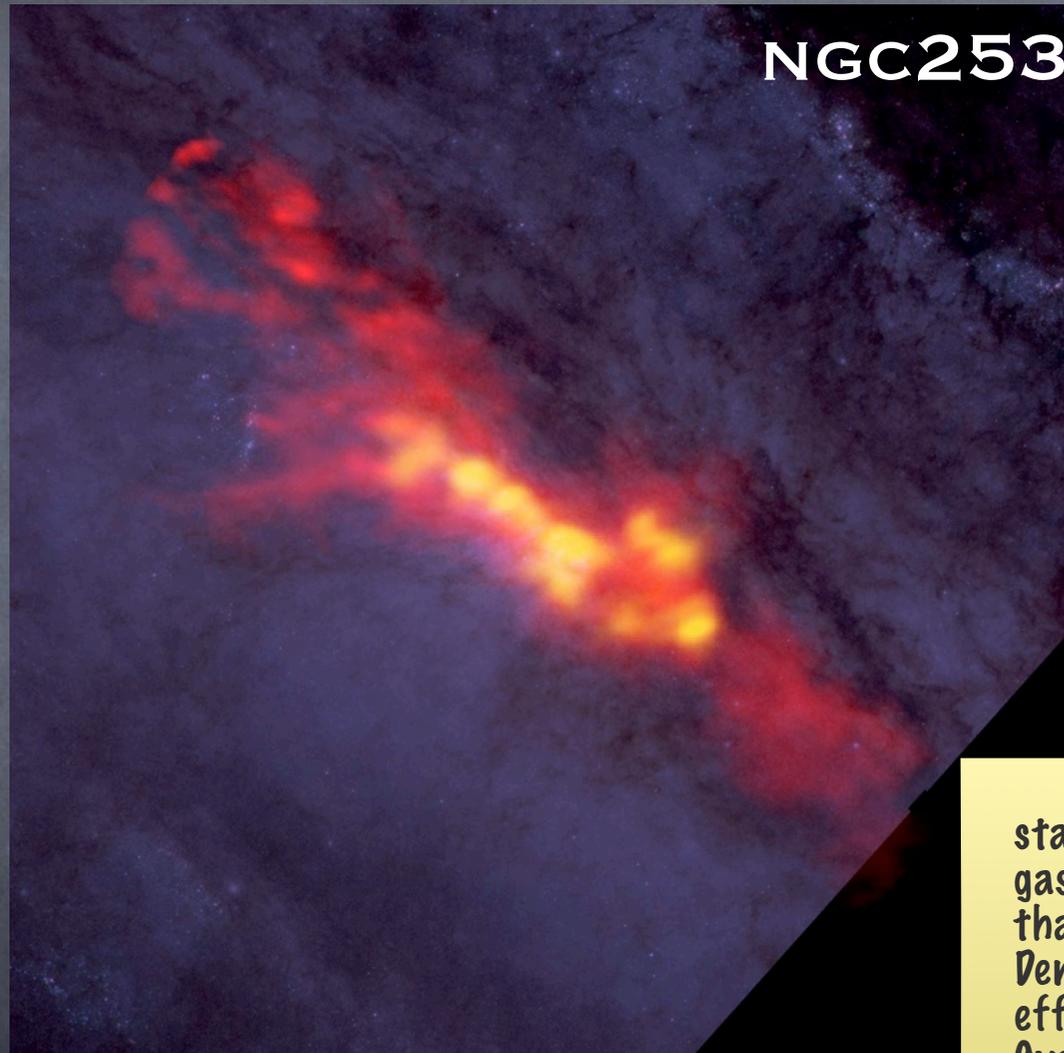
# LOCAL UNIVERSE



HERRERA ET AL. (2012, FIG.1)

- spatial resolution
- comparison to H<sub>2</sub>
- SciVerif

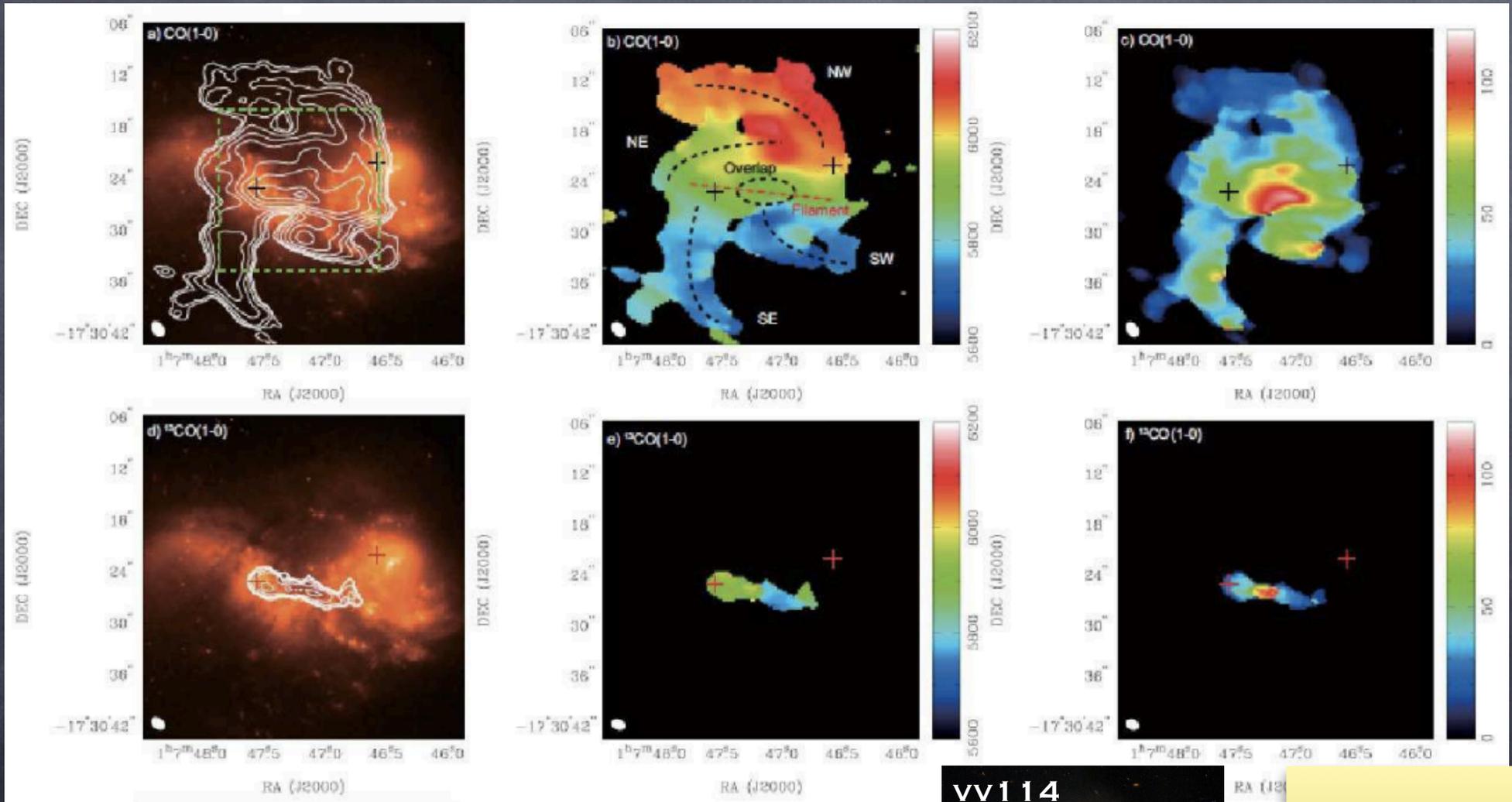
# LOCAL UNIVERSE



LEROY ET AL. (2015, PRESS RELEASE)

star formation efficiency  
gas consumption 10x faster  
than in disc  
Density is key setting  
efficiency  
Cyc0&1

# LOCAL UNIVERSE

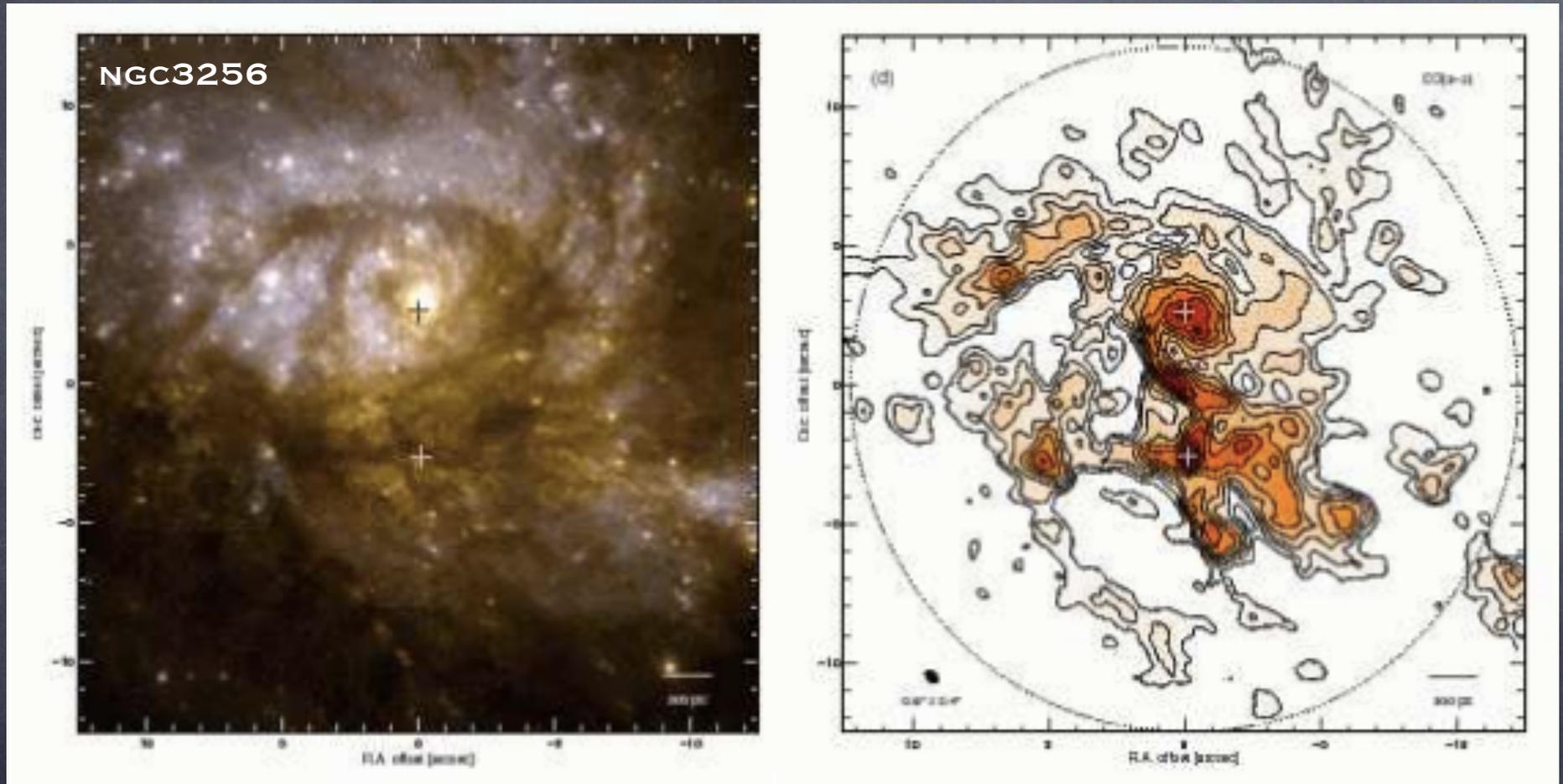


SAITO ET AL. (2015, FIG.2)



- dynamics
- tidal tail
- isotope
- 40min Cyc0

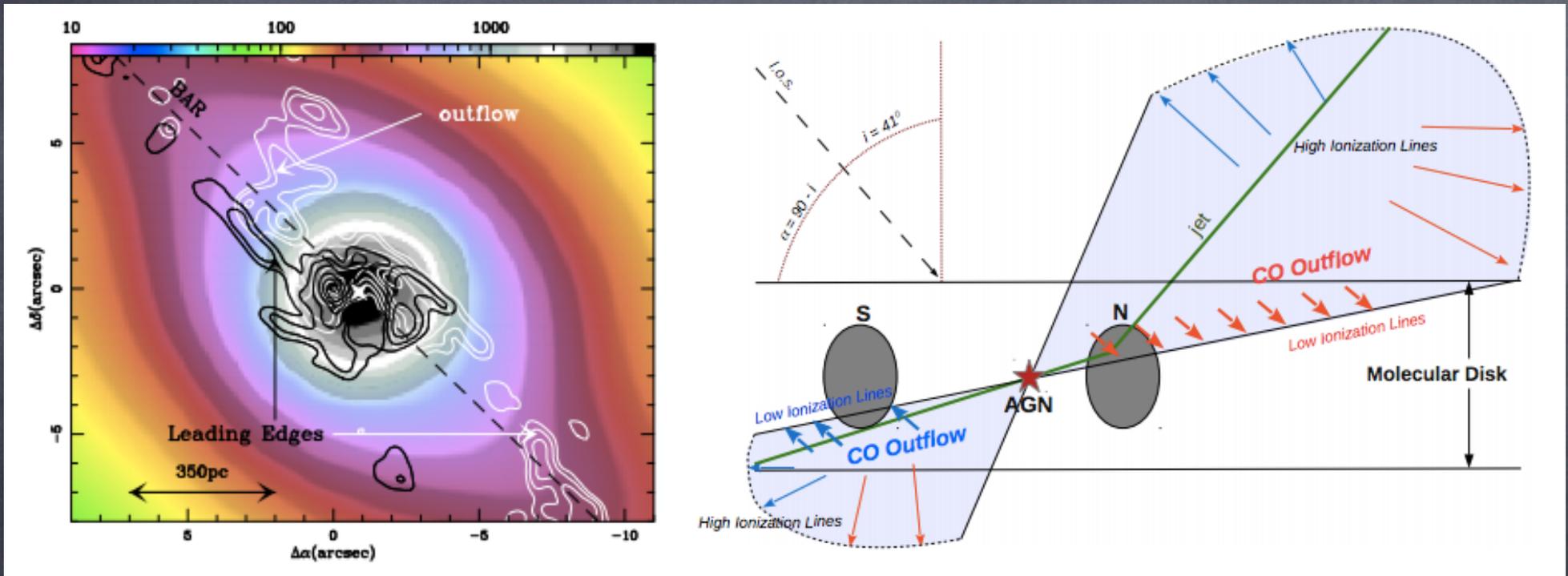
# LOCAL UNIVERSE



SAKAMOTO ET AL. (2014, FIG.18)

- outflows:  
star-formation  
plus agn  
- sciVer+CO, B7,  
2x3h

# LOCAL UNIVERSE



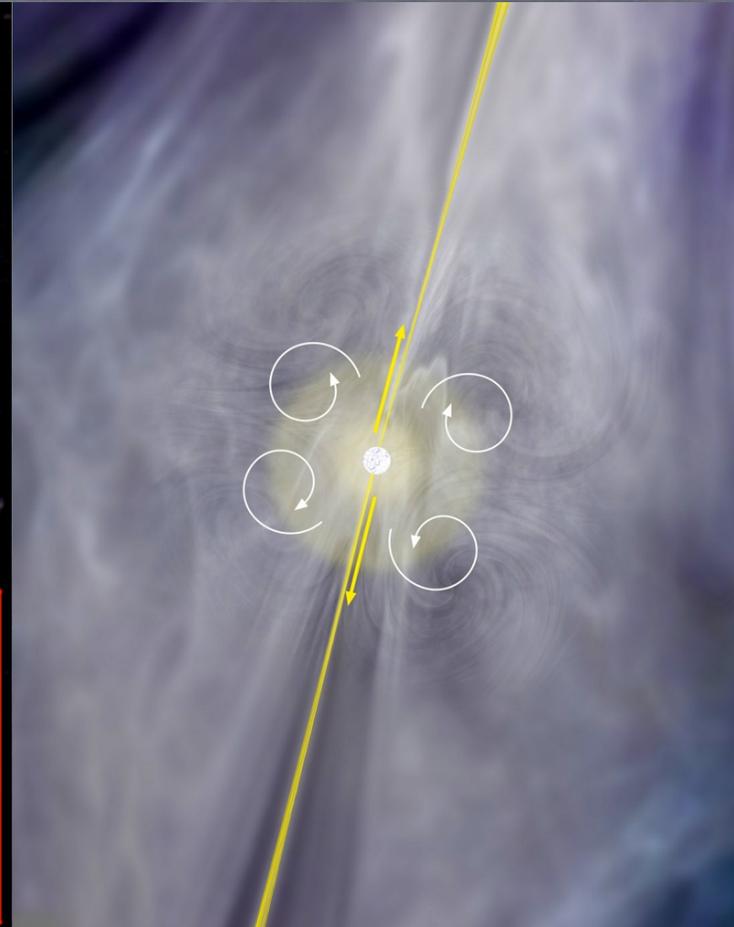
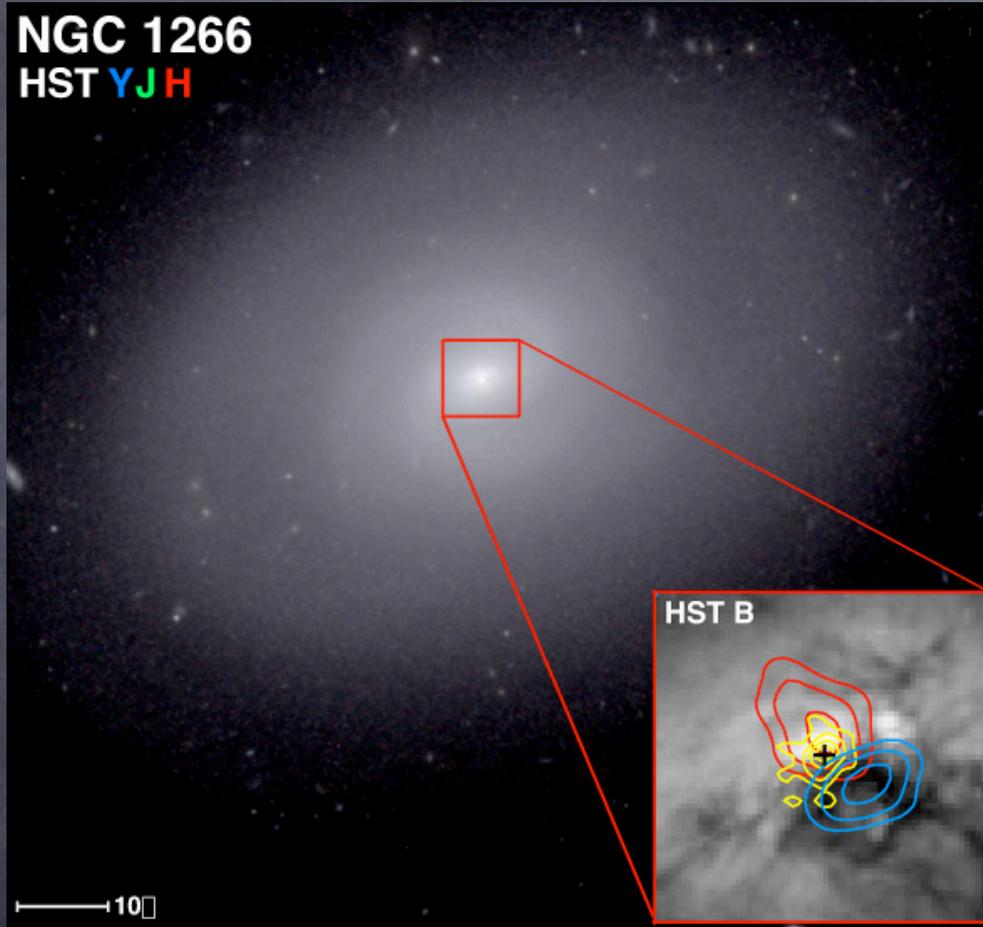
GARCÍA-BURILLO ET AL. 2014, FIGS. 5 AND 17)

NGC 1068



outflows  
 - B7+9, Cyc0  
 - 30min/track  
 (138tot)

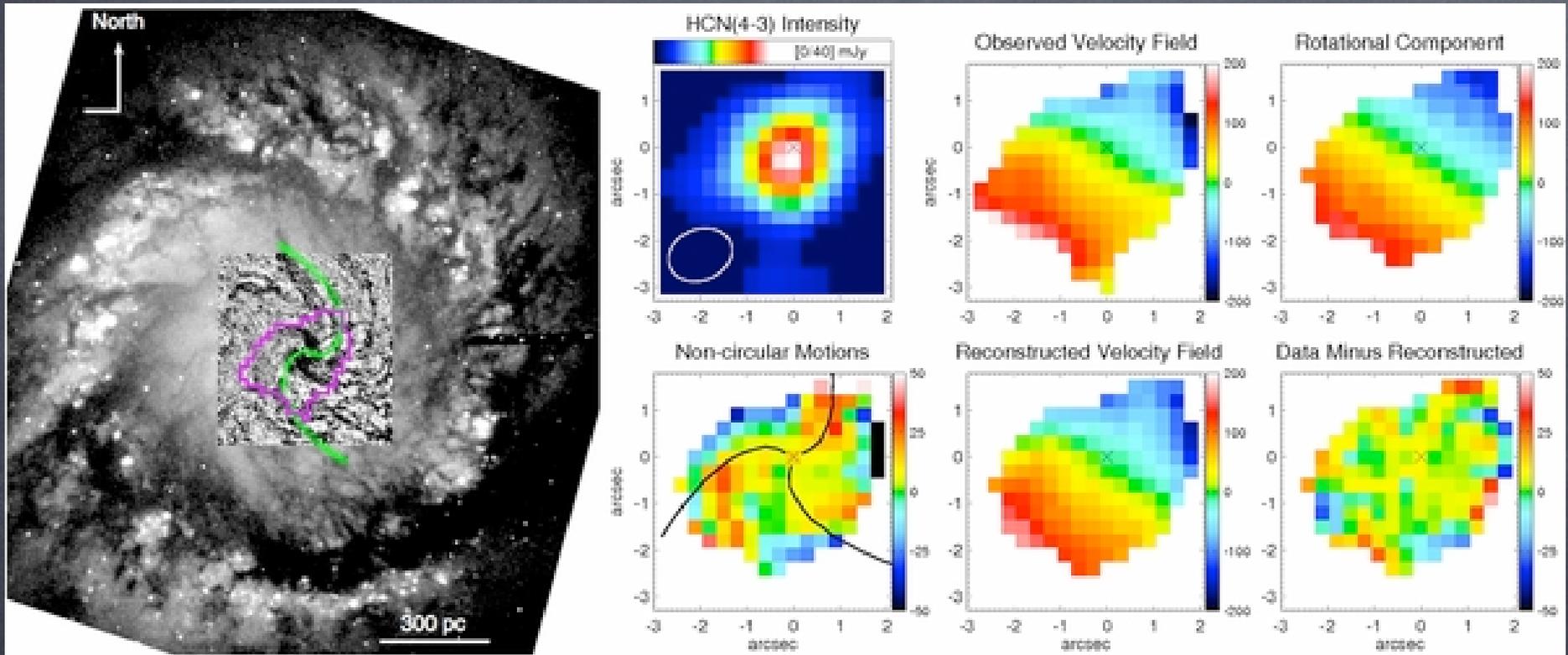
# LOCAL UNIVERSE



ALATALO ET AL. (2015, PRESS RELEASE)

outflows  
also CARMA  
too much gas for the SFR  
too much turbulence  
Cyc0

# LOCAL UNIVERSE

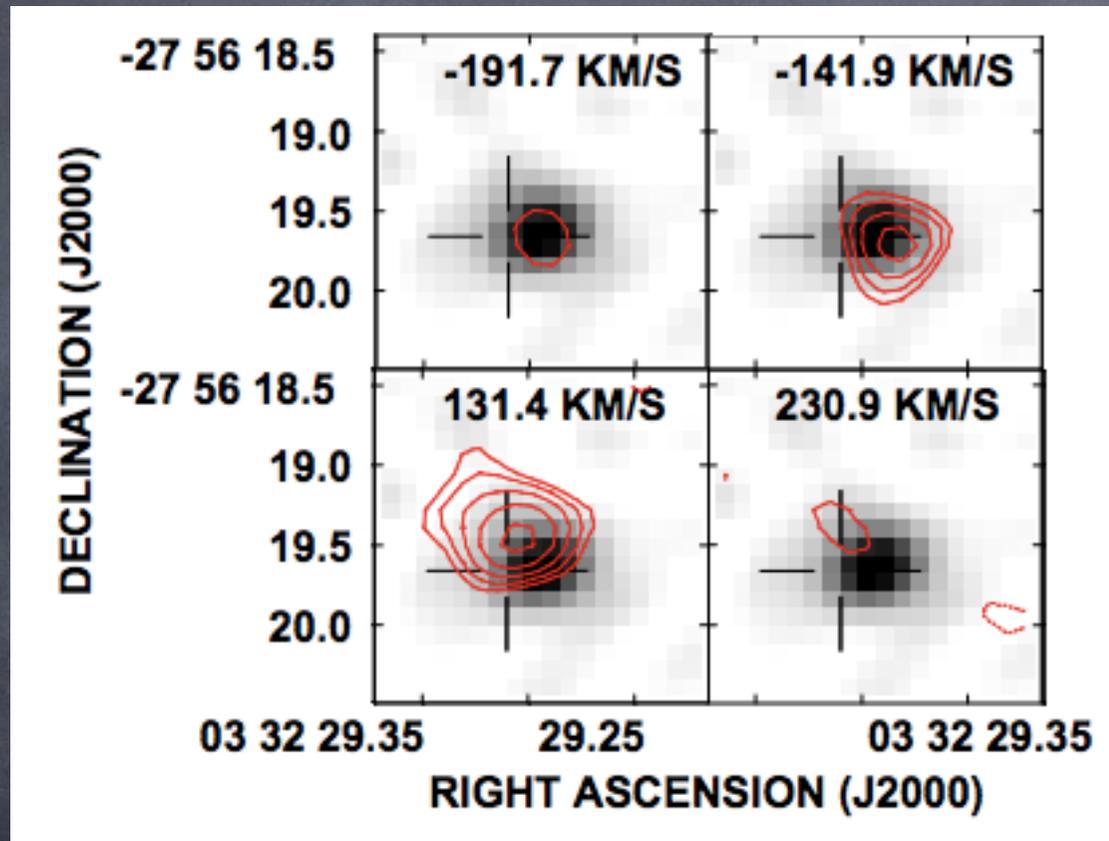


FATHI ET AL. (2013, FIG. 1)

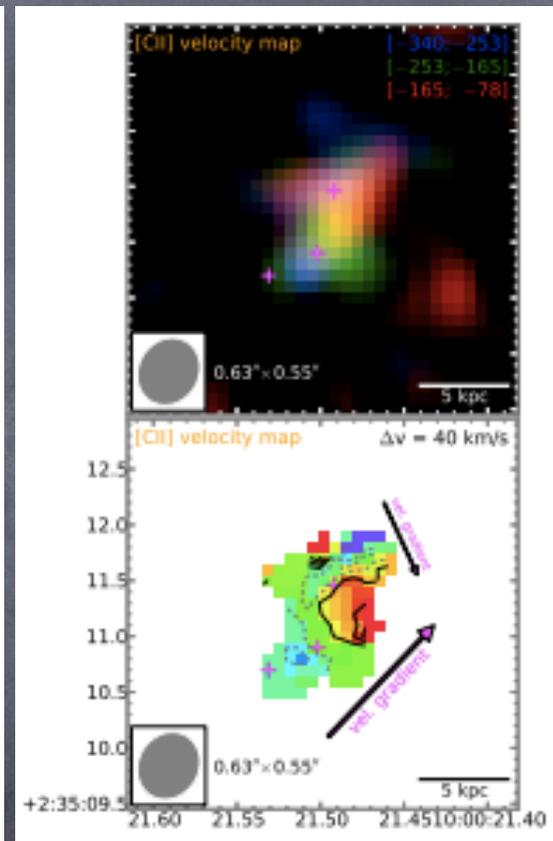
[WAIT ALSO FOR TREMBLAY ET AL. (ABELL 2597)]

inflow gas  
Cyc0

# HIGH-Z UNIVERSE



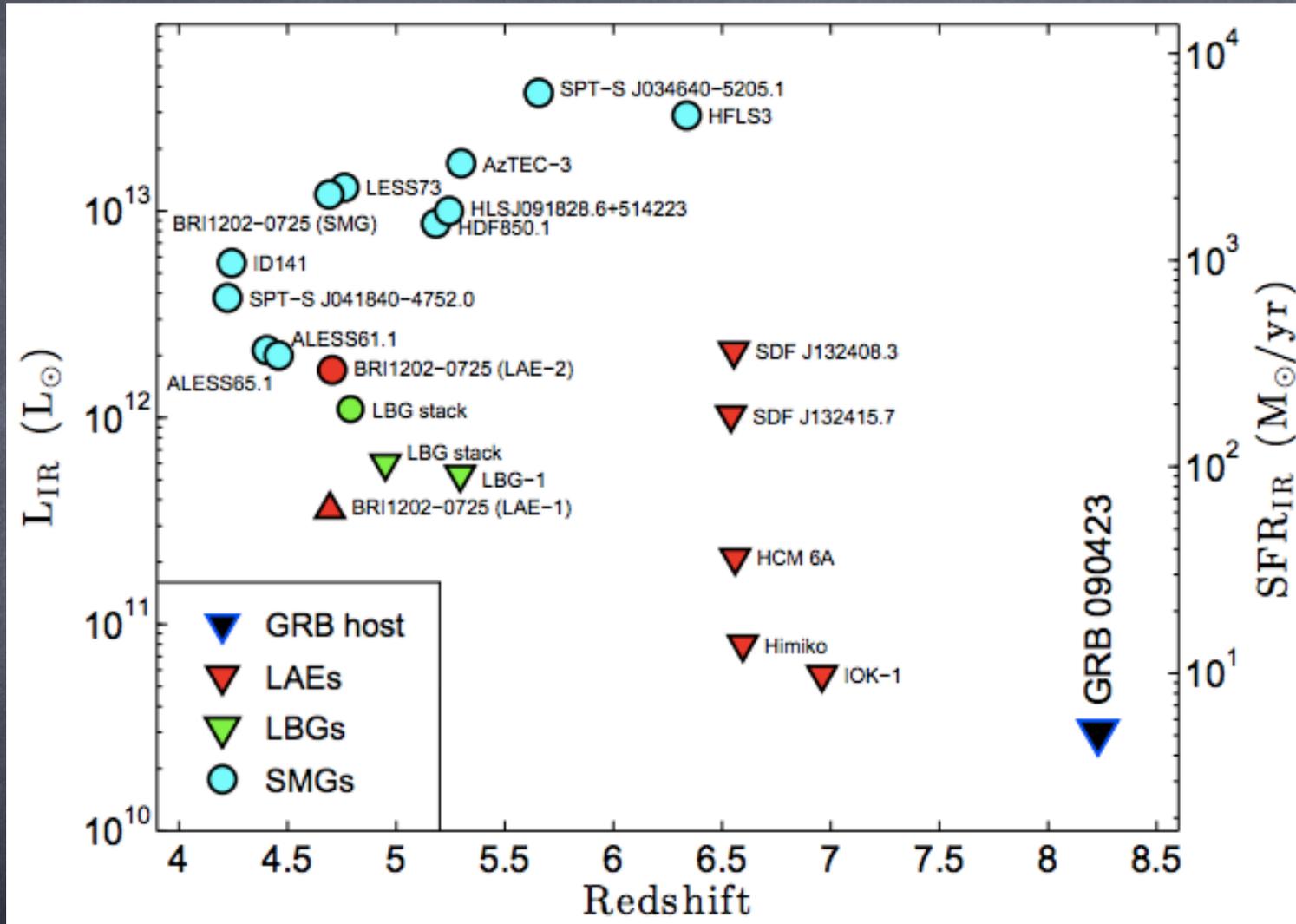
DE BREUCK ET AL. (2014, FIG. 3)



RIECHERS ET AL. (2014, FIG. 7)

cii dynamics,  $z \sim 4-5$   
 Cyc0, B7  
 Riechers, 125min

# HIGH-Z UNIVERSE

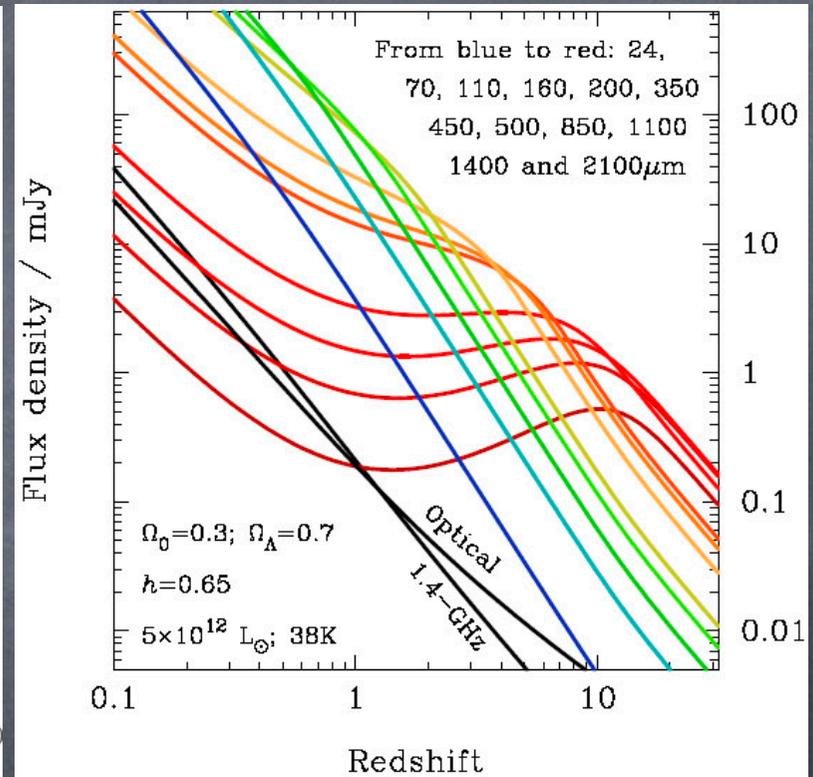
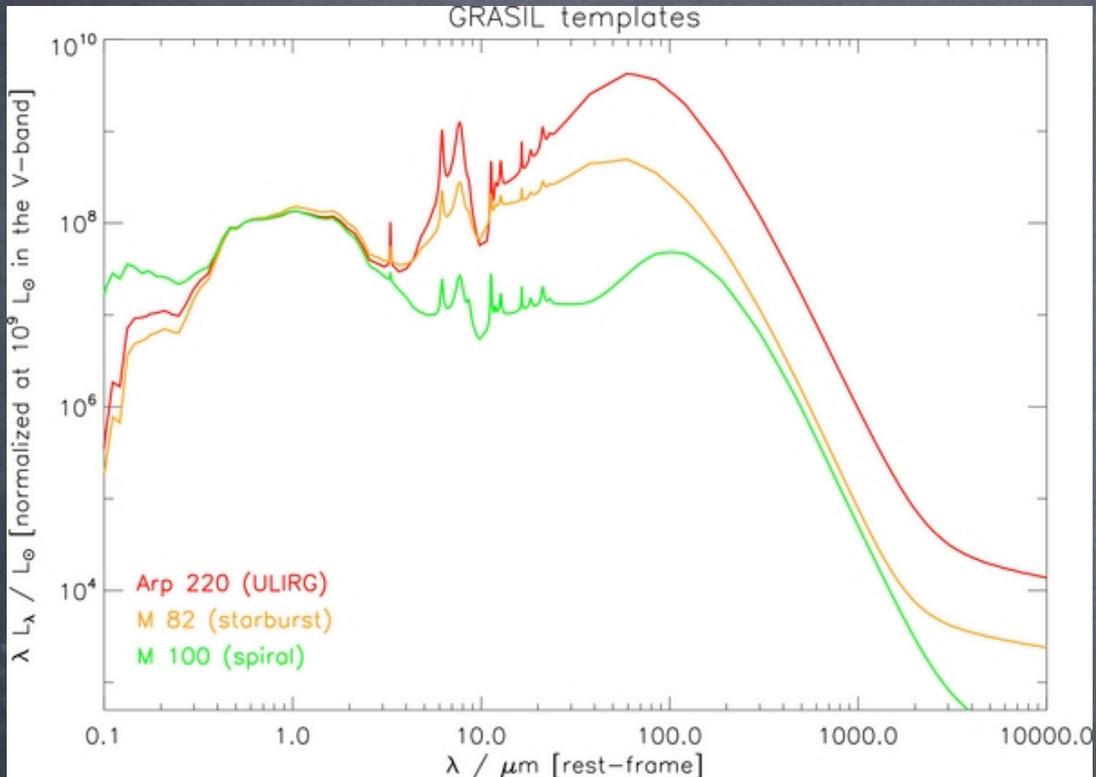


BERGER ET AL. (2014, FIG. 3)

[BUT WAIT FOR D. WATSON ET AL. (2015, NATURE, A1689-ZD1, z=7.5)]

high-z continuum  
 detection  
 GRB Cycl. B6, 163min

# HIGH-Z UNIVERSE



DA CUNHA ET AL. (2013, FIG. 12)

BLAIN ET AL. (2002, FIG. 4)

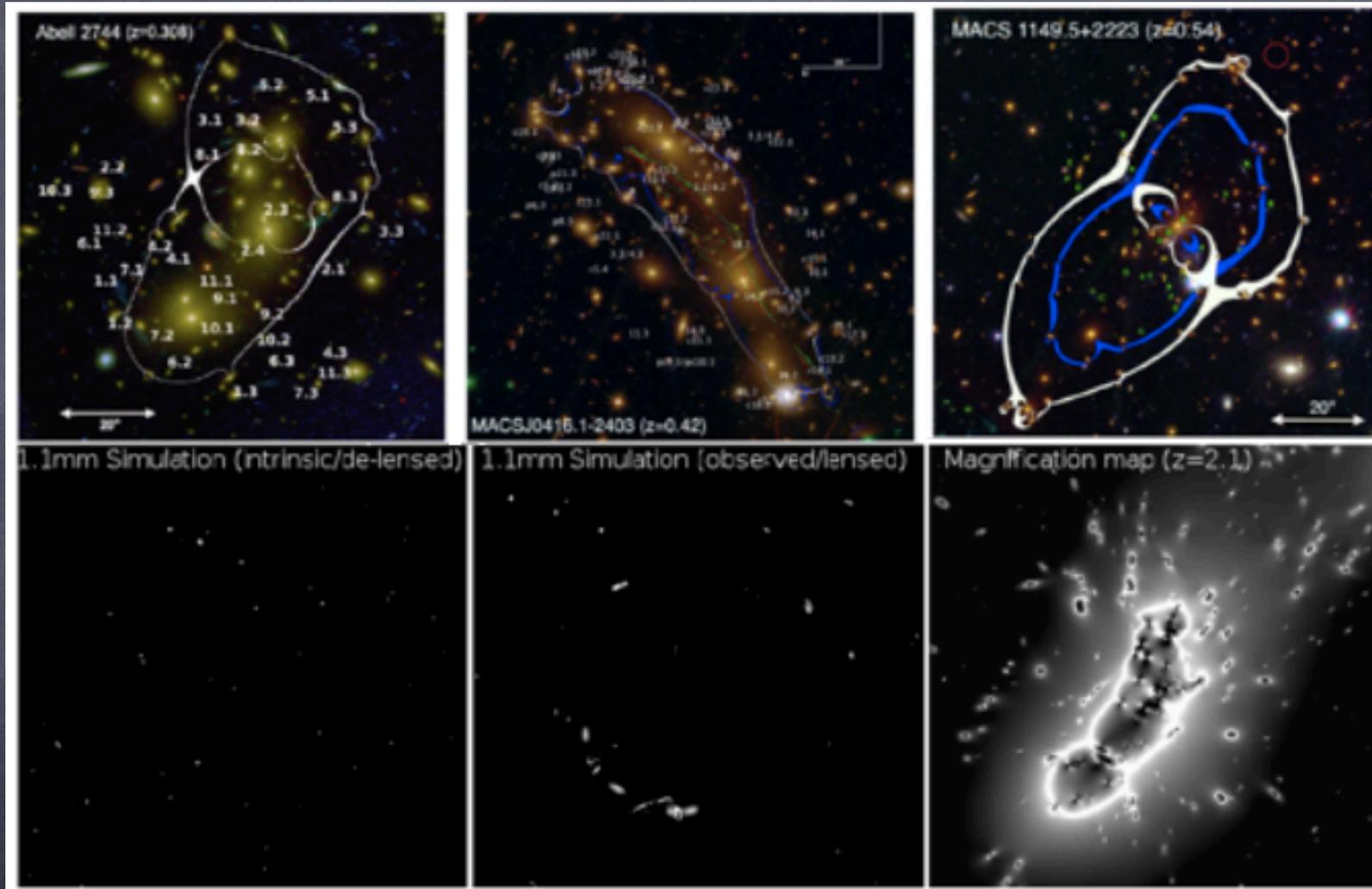
k-correction  
 mm high-z bias

# HIGH-Z UNIVERSE

- AN ALMA 1.3-MM IMAGE OF THE HUBBLE ULTRA DEEP FIELD  
[PI JAMES DUNLOP]
- MINING AND RESOLVING A CONTIGUOUS 1.5 ARCMIN<sup>2</sup> WINDOW IN THE SXDF-CANDELS-UDS AT 1.1 MM WITH ALMA  
[PI KOTARO KOHNO]
- A MOLECULAR ALMA DEEP FIELD IN THE UDF  
[PI FABIAN WALTER]
- UNVEILING THE POPULATION OF HIGH-REDSHIFT SUBMILLIMETER GALAXIES WITH ALMA 1.2 MM IMAGING  
[PI MANUEL ARAVENA]
- AN ALMA 1.3 MM SPECTROSCOPIC SURVEY IN THE HUBBLE ULTRA DEEP FIELD  
[PI MANUEL ARAVENA]
- LENSING THROUGH COSMIC TIME: ALMA CONSTRAINTS ON "NORMAL" GALAXIES IN THE HST FRONTIER FIELDS  
[PI FRANZ BAUER]
- EVOLUTION OF ISM IN STAR-FORMING GALAXIES AT  $z = 1 - 5$   
[PI NICK SCOVILLE]

many deep fields  
mm bias high-z  
goal continuum detect  
normal galaxies

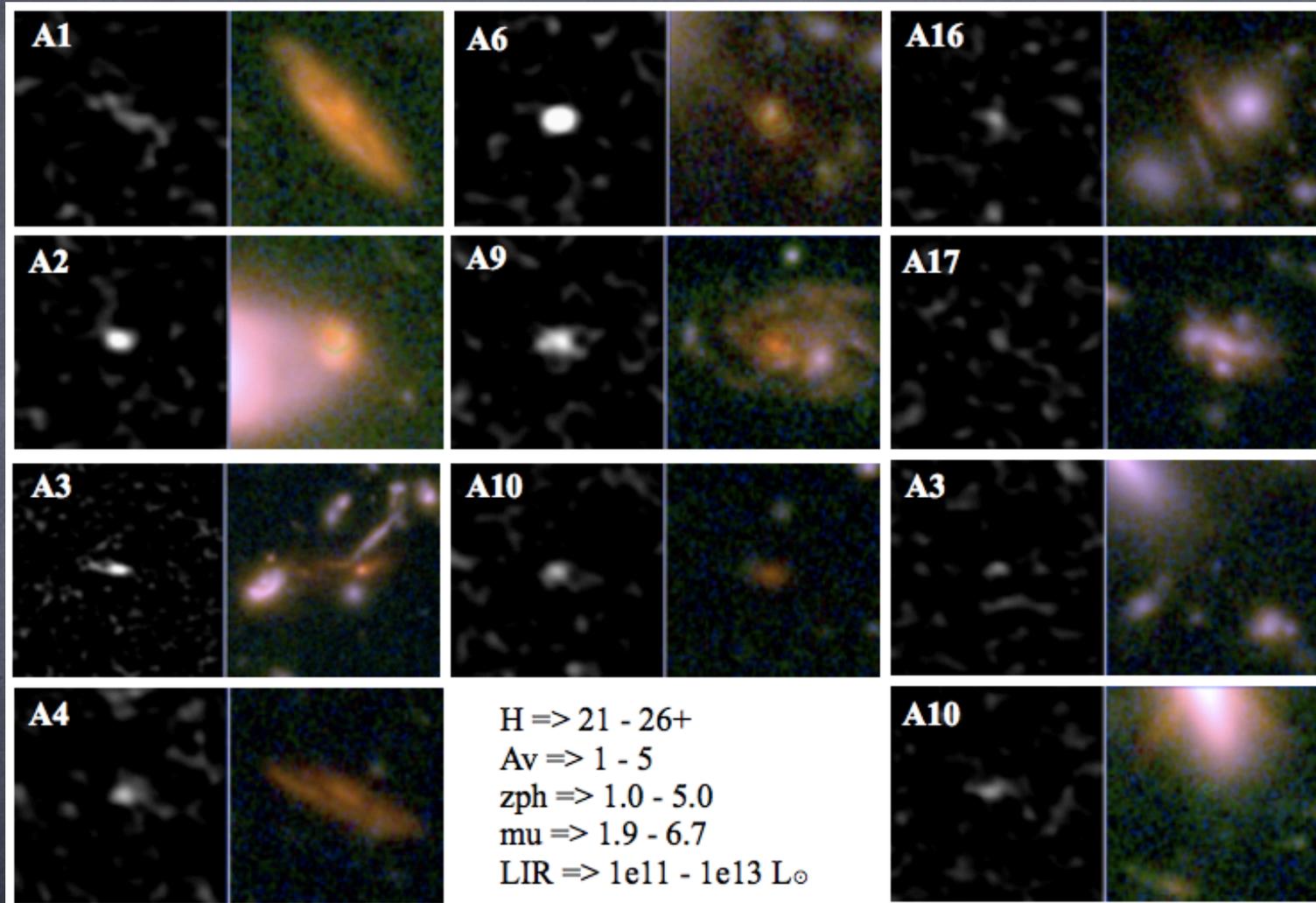
# HIGH-Z UNIVERSE



PI F. BAUER — LENSING THROUGH COSMIC TIME

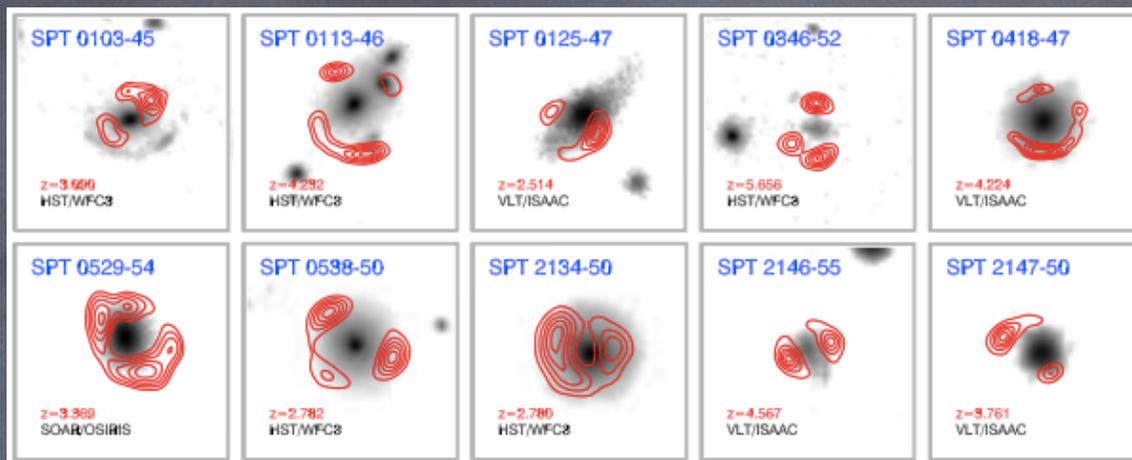
k-corr + lensing

# HIGH-Z UNIVERSE

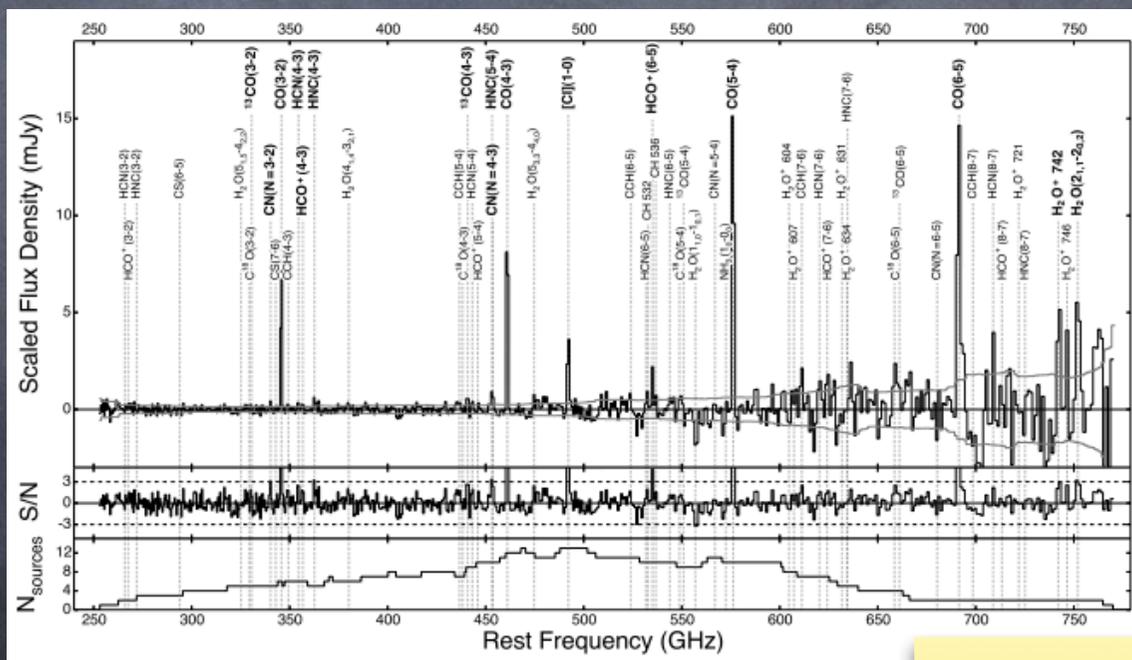


PI F. BAUER — LENSING THROUGH COSMIC TIME

# HIGH-Z UNIVERSE



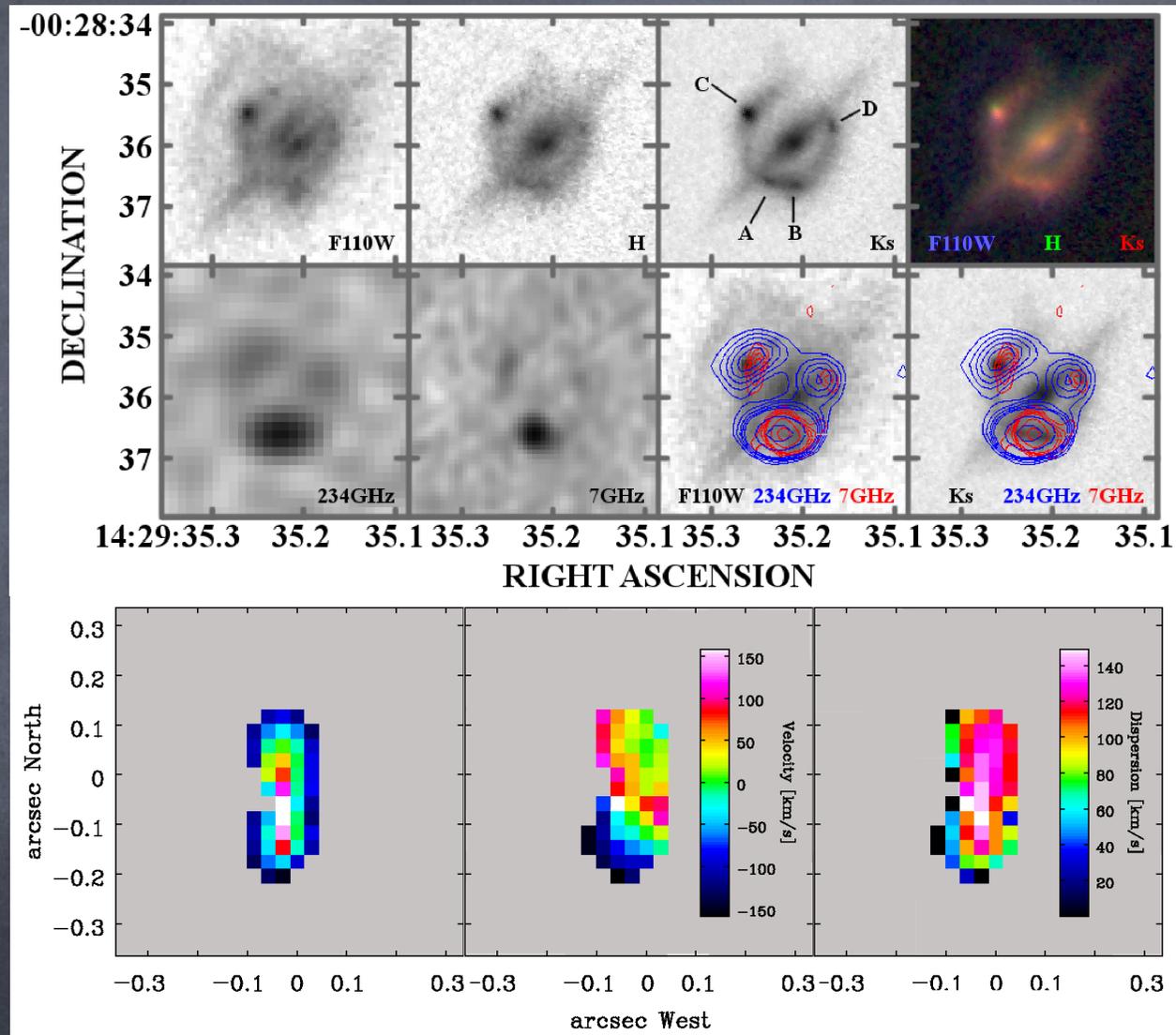
VIEIRA ET AL. (2013, FIG. 1)



SPIPKER ET AL. (2014, FIG. 2)

spt sample  
 spec detection  
 Cyc0, ~12min

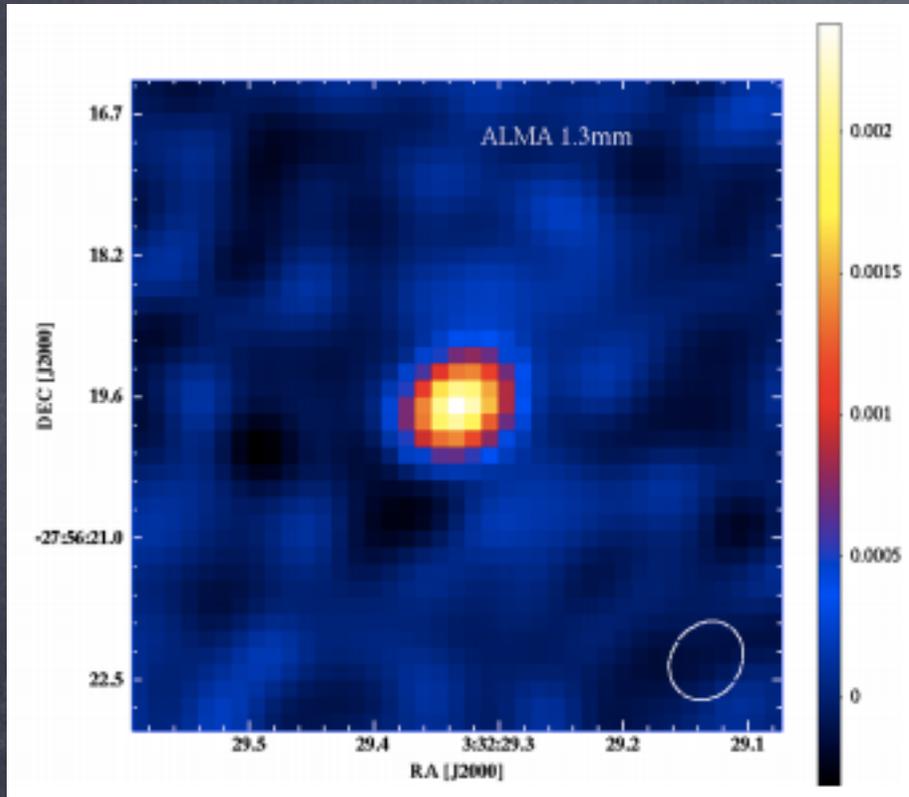
# HIGH-Z UNIVERSE



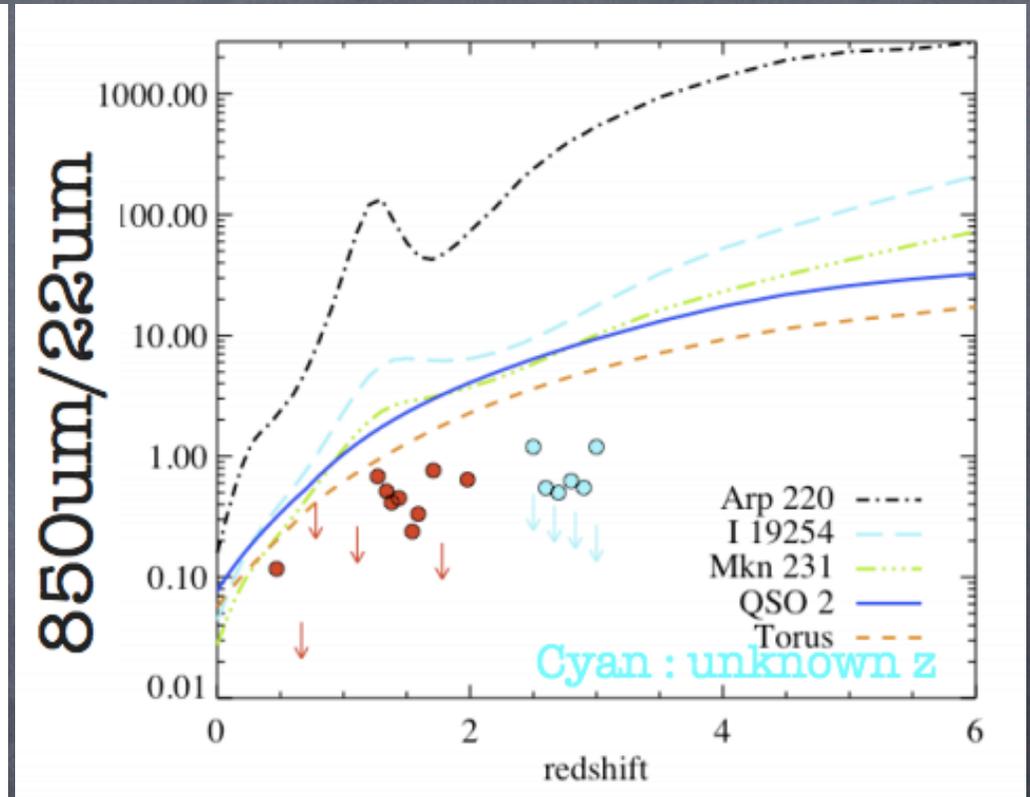
MESSIAS ET AL. (2014B, FIG. 1 AND 10)

lensed dynamics  
Cyc0, 30min

# HIGH-Z UNIVERSE



GILLI ET AL. (2014, FIG. 1)

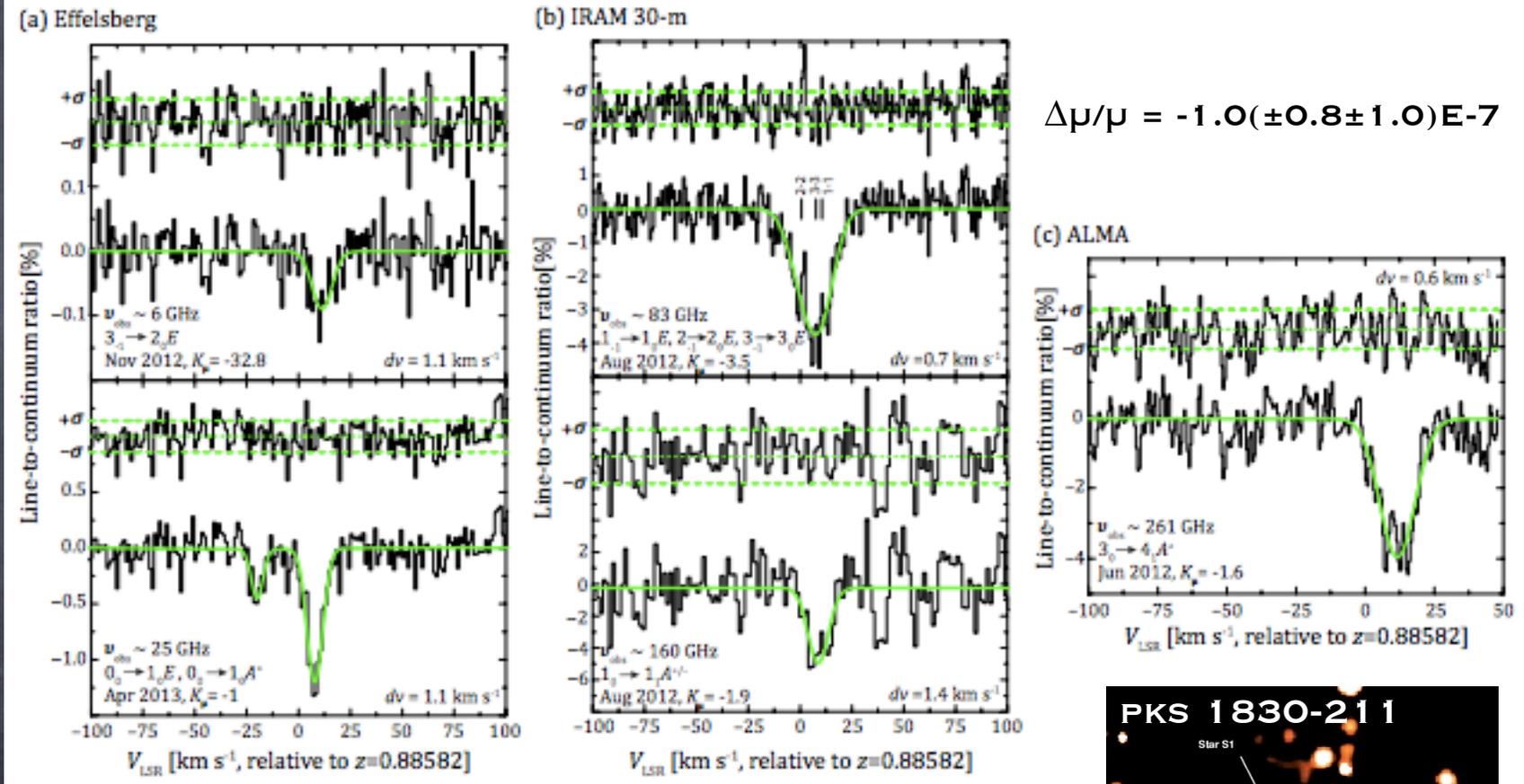


PI C. LONSDALE — NVSS-WISE AGN

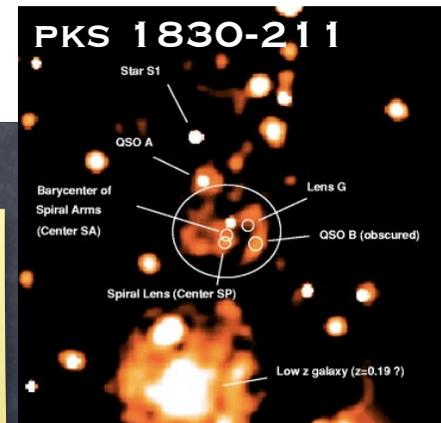
## agn hosts

- G14 host column density = x-ray  $N_h$ , no need pc-absorber
- general agn dominance (c.lonsdale)
- Cyc0, B6, 3min (Gilli)

# COSMOLOGY

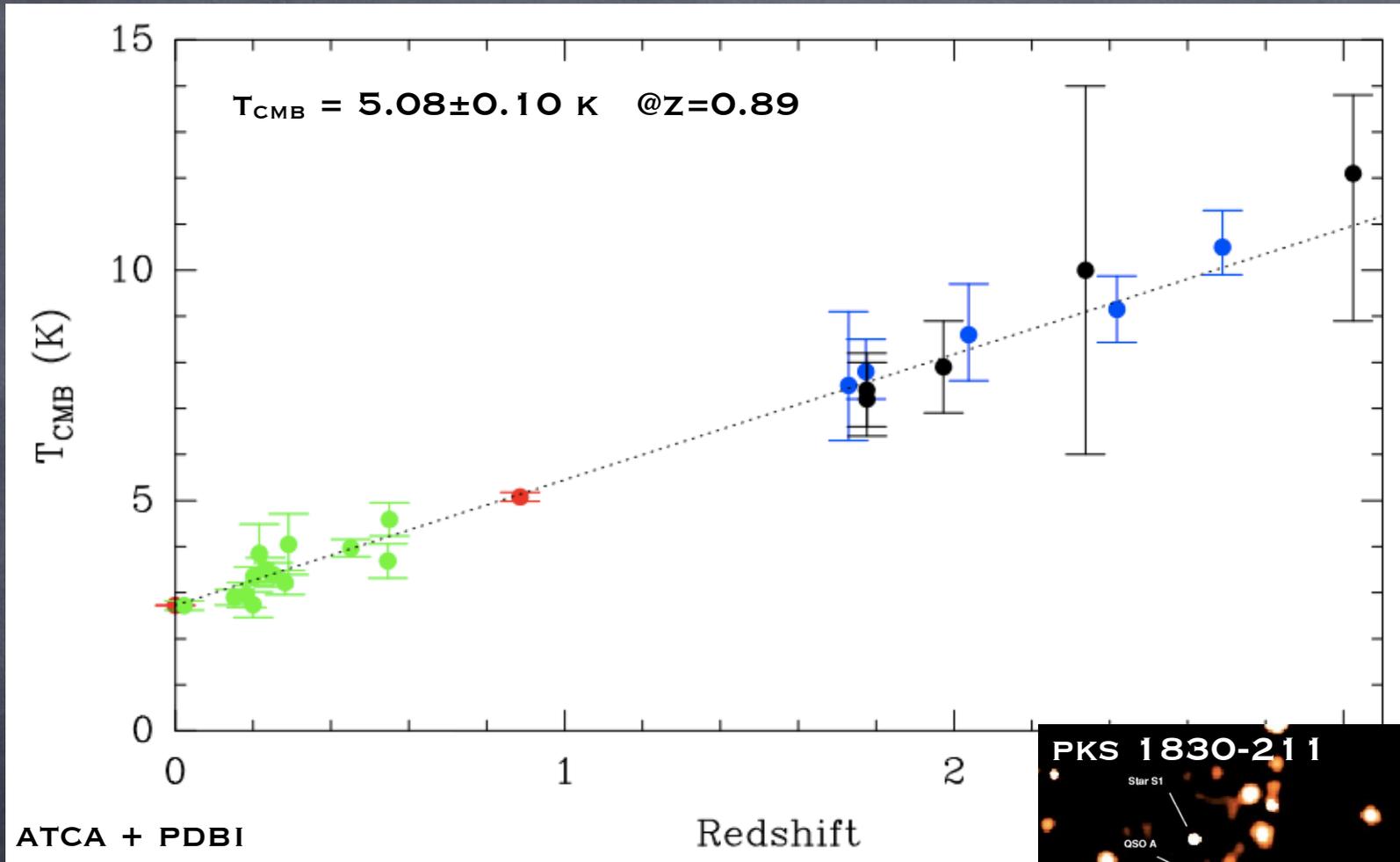


BAGDONAITE ET AL. (2013, FIG. 1)

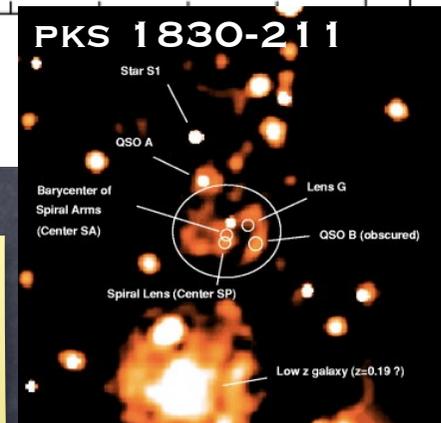


proton-electron mass  
 ratio toward pks1830  
 Null result  
 Cyc0

# COSMOLOGY

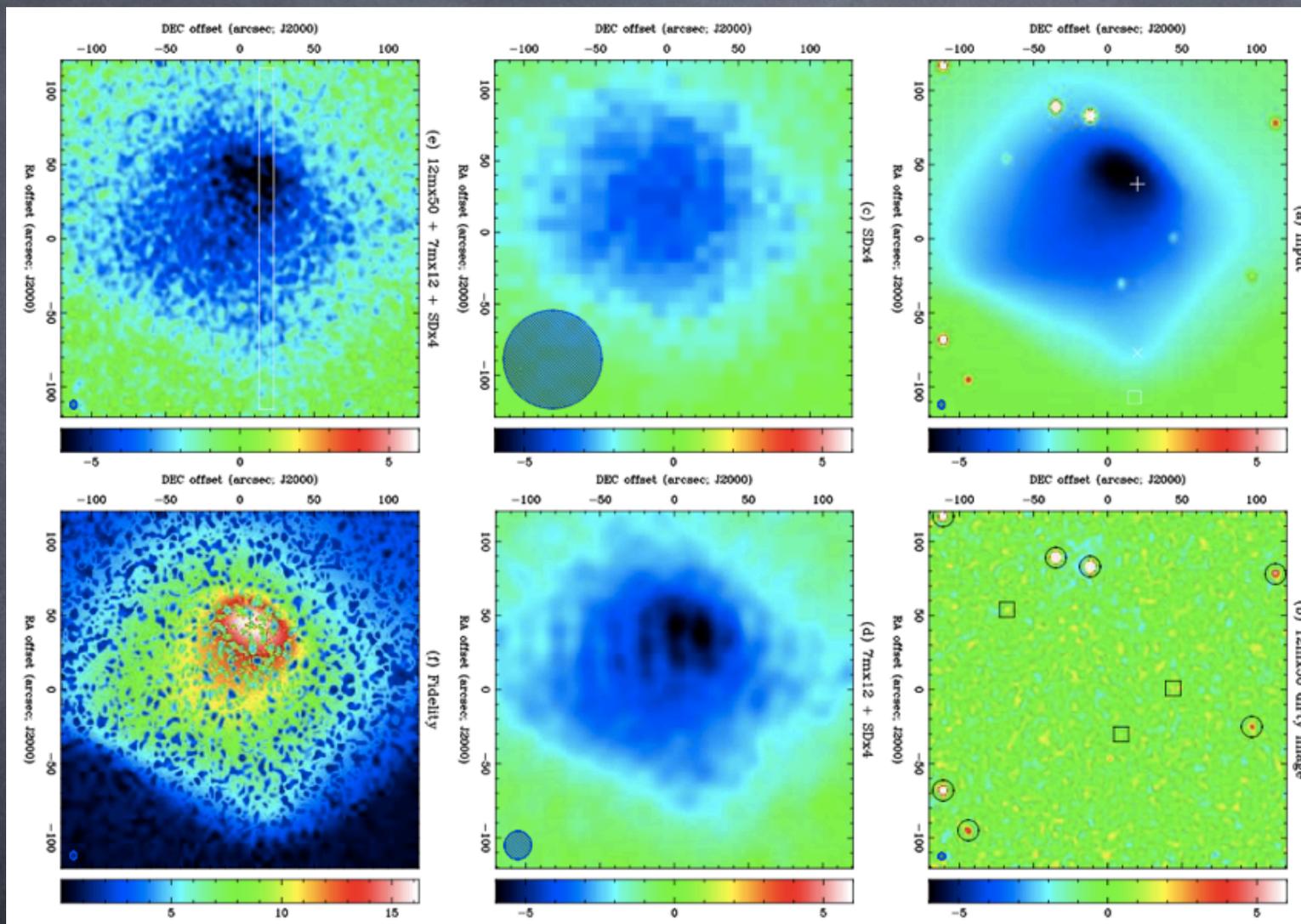


MULLER ET AL. (2012, FIG. 5)



t<sub>cmb</sub> toward pks1830  
 adiabatic expansion  
 atca+pdbi

# COSMOLOGY



YAMADA, KITAYAMA, ET AL. (2012, FIG. 5, SIMULATION)

thermal sz effect  
10h(12m), 40h(7m,TP)

THANK YOU!